Abstract:

Line 1: with atmospheric → using atmospheric

Line 3: by the air showers. → by charged particles belonging to the air shower generated by the primary cosmic ray.

Line 3: this background → this background effect

Introduction

The first two sentences (lines 13-18) are slightly convoluted. I suggest modifying for instance in the following way:

"Two aspects of muon tomography of volcanoes make relevant, if not necessary, the use of Monte-Carlo simulations to accurately evaluate the attenuation of the flux of atmospheric muons crossing the volcano. The first aspect is related to the fact that the muon flux has to be described with sufficient precision and analytical approximations cannot easily be used, in particular at low muon energies and for nearly horizontal muons."

Line 23: remove "such"

Line 25: Question on "Various particles produced in air showers can hit the detector coherently". In the following of the paper, only muons are considered in the Monte-Carlo as "background" signal. However, those atmospheric muons are the "signal", whose enhancement/attenuation gives the muon tomography. I expected that the background can be due also by other charged particles in the air shower (namely, electrons and positrons) faking a muon signal. Do you plan to insert these in the MC code in the future? Could you comment?

Section 2

Line 17: "Doing so, the most energetic component of their flux initiates events called" → High energy cosmic rays initiates

Line 23: photons or leptons → photons and leptons

Line 10 (decay of mesons): write 2.6, 1.2 and 5.1 instead of 2,6 1,2 and 5,1

Section 3

Line 11: Such a code → The full program chain

Line 14: The setup of the simulation at present day \rightarrow The present simulation setup

Line 18: The physics list used → The used interaction model

Question on Fig. 1. According to other simulations, the average number of secondary muons per shower is much higher. For instance, from Fig. 8.11 of the Stanev book (High energy cosmic rays), about 100 muons with 1 GeV threshold and 10 muons with 30 GeV threshold are expected from a 10 TeV primary proton at sea level. If I integrate the solid line of Fig. 1, about 1 muon is obtained. Do the lines correspond to the average number of muons PER mesons? I suppose that the difference of about 2 orders of magnitude is not due to the 870 m. level.

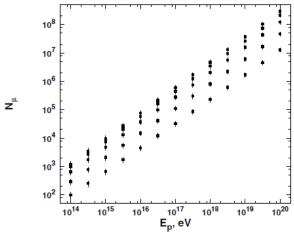


Fig. 8.11. Average number of muons at sea level in vertical proton showers as a function of the primary energy in GeV. From top to bottom the muon energy thresholds are 0.3, 1, 3, 10, and 30 GeV.

Line 4 of an inclusive cosmic ray flux \rightarrow of a cosmic ray flux which includes all the nuclear species

Section 4

line 15: "realistic"?? You mean with a power law with the correct spectral index?

line 19: Question: why do not you restrict only to proton primary? This could have an effect on the discrepancies on time distributions?

line 23: and temporal distributions \rightarrow and arrival time distribution

line 5: towards wider high energy showers. \rightarrow towards wider showers at high energies.

Conclusions

line 14: in GEANT4 → in the framework of the GEANT4 simulation code

line 17: that we obtained with vertical → produced by vertical

line 18: a well-tested air shower simulation software: CORSIKA \rightarrow the well-tested CORSIKA air shower simulation code.

line 23: again "an inclusive", which is not clear. See above