

1 **Reply to Referees**

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3 **Ryuichi Nishiyama on behalf of the authors.**

4 **19<sup>th</sup> Feb. 2014.**

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6 Thank you for your careful reading, comments and language corrections. Followings are my  
7 replies to your comments. This discussion paper consists of two parts. In the former part, we  
8 reply to the comments given to some topics from several referees. In the latter part, we  
9 answer to the specific comments by each referee.

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12 **English problems**

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14 It would certainly benefit from a revision by a native English speaker. Expressions  
15 used throughout the text are often incorrect and sentence construction belies very  
16 little confidence with the language used (**Referee #1**).

17  
18 I feel that English language should be improved throughout the manuscript. In the  
19 following, I give some advices on how to improve some sentences. Anyway, I am  
20 not a mother tongue and I urge the author to have the manuscript checked by one  
21 of them for linguistic correctness (**Referee #4**).

22  
23 Responding to your comments, I already asked a native English speaker for correcting our  
24 manuscript. I think linguistic correctness becomes acceptable in the next revision.

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26  
27 **Remarks on muography detectors other than emulsion films**

28  
29 Introduction (1) As stated earlier, it would benefit from a brief overview of what type  
30 of detectors have been used so far in this application (**Referee #1**).

31  
32 Furthermore, the authors should provide more detailed discussion on the  
33 comparison between the performances of emulsion film detectors and other types  
34 of detectors like, for example, those employing plastic scintillators (**Referee #4**).

35  
36 To address this point, a more detailed discussion should be provided about

37 advantages and limits of emulsion film detectors, compared with detectors based  
38 on different principles (**Referee #4**).

39

40 There are several types of detectors such as scintillation type (Tanaka et al., 2007; Lesparre  
41 et al., 2010; Anastasio et al., 2013, etc), gas chamber type (Cârloganu et al., 2013; Barnafoldi  
42 et al., 2013, etc) for the application of muon radiography. I will address these precedent works  
43 in the next revision. Compared with these types of detectors, the greatest advantages of  
44 emulsion detectors are their portability and very high resolution (position resolution: a few  
45 microns, angular resolution: a few milli-radians). This work benefited greatly from the high  
46 angular resolution, which enables momentum selection. One disadvantage of emulsion is  
47 that it does not provide time information of particle arrival. The comparison of the  
48 performances should be discussed after we install several detectors at same site. This will  
49 be a subject of future study. We will address these issues in the next revision.

50

51

## 52 **Grain density cuts**

53

54 Section 5.2.1 The first sentence is incomprehensible. I surmise that the authors cut  
55 away events with tracks having grain densities higher than a certain threshold. If  
56 this interpretation is correct, the text should be modified accordingly (**Referee #1**).

57

58 I would suggest replacing “an average grain density (number of AgBr grains on  
59 track) higher than threshold. The grain density is” by “grain densities higher than a  
60 certain threshold. The grain density is number of silver grains per unit length along  
61 the track and is” (**Referee #2**).

62

63 P656: -I 5: Please rewrite the first sentence which looks strange (**Referee #3**).

64

65 The first sentence in Sec. 5.2.1. in the previous manuscript was strange and misleading. I  
66 follow Referee #2's suggestion. Thank you for your careful reading.

67

68

## 69 **Need for more description on experimental conditions**

70

71 Fig. 7: How did the authors estimate the shaded histograms in Fig. 7? Why do the  
72 tracks recorded during transportation have lower numbers of hits? Were the

73 detectors fabricated just before the installation? Better to clarify experimental  
74 conditions (**Referee #2**).

75

76 The detectors were fabricated just before the installation. Before the installation, all the  
77 OPERA film were refreshed in high temperature and humidity place. We estimate the amount  
78 of fake signals by giving misalignment of 500 microns between each film and selected tracks  
79 using the same method as employed in the true signal selection. I will address this in the next  
80 revision.

81

82 Readers also would like to have an idea of the data taking conditions: environmental  
83 parameters effect, humidity/temperature controls in any, fog density in the  
84 emulsions etc (**Referee #3**).

85

86 The observation was taken in place in a warehouse without air-conditioning in the midst of  
87 winter season in northern part of Japan. We did not monitor temperature or humidity during  
88 exposure. According to the Japan Meteorological Agency's report, the maximum temperature  
89 is 15.9 degrees in Celsius, and the minimum is – 8.8 degrees around the detector site during  
90 the exposure period. Although we did not monitor the humidity, it did not matter to the  
91 performance of our detectors, because the films were packed inside of the envelopes. We  
92 will try to give specific conditions as much as possible in the next revision.

93

94

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95 **Reply to each referee's comments**

96

97 **To Referee #1**

98

99 "Thus one is left with a paper which in reality describes an improvement  
100 (significant) to a standard apparatus used for muon radiography.

101

102 I do not get the meaning of this sentence. Could you please give us some explanations if our  
103 answers below are not satisfactory to you ?

104

105 I would thus change the title of the paper to reflect the focus on detector  
106 improvement and suggest that the authors further develop this aspect in their  
107 conclusions, covering feasibility of large area detectors and deployability in the field.

108

109 The most important characteristic of emulsion film for muography is that it is very easy to  
110 install in the field. The possibility of large area detectors should be discussed along with the  
111 current improvement of read-out system (eg. Morishima & Nakano, 2010). I will address  
112 these topics in the next revision. However, I do not want to change the title of our paper,  
113 because this work focuses on the background source of muography and does not deal with  
114 technical aspects of emulsion film itself. In response to your suggestions, we have an  
115 alternative title of our paper, which is “Experimental study of momentum of background  
116 particles in muon radiography using emulsion film detectors”.

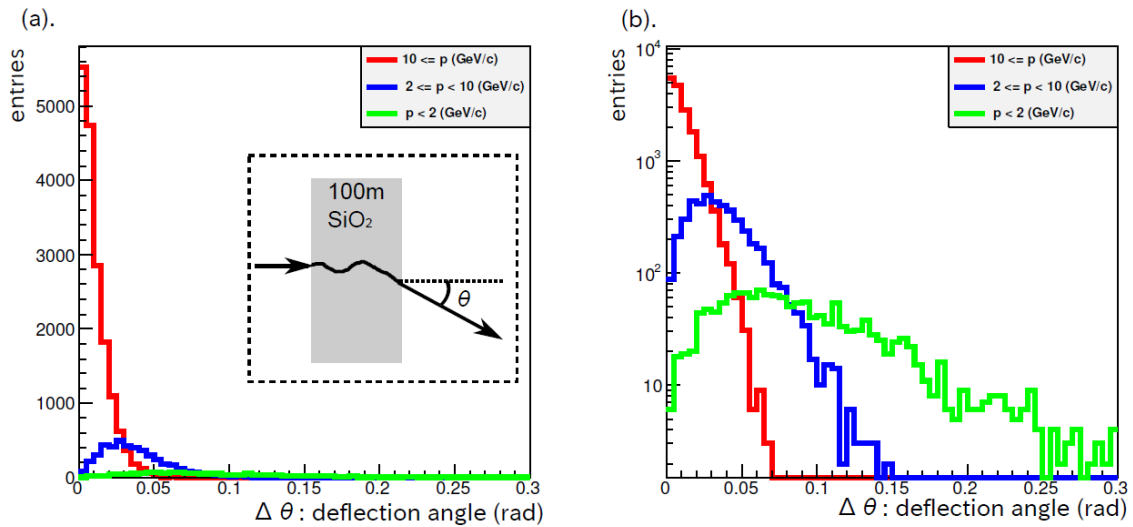
117

118           Also it’s not clear why the authors suppose straight away that backgrounds to muon  
119 radiography are due only to low momenta particles. Could they also be due to high  
120 momentum muons coming from behind (scattered on the ground surface) ?

121

122 In this paper, we do not exclude the possibility that muons coming from behind the detector  
123 could cause background noise. However, we want to emphasize that such muons, which  
124 were scattered in the ground are expected to be not so energetic. Please see a revised  
125 Figure 1 (below). This histogram shows the distribution of deflection angles of muons after  
126 passing through a certain thickness of rock. This figure shows that muons with momenta  $>$   
127  $10 \text{ GeV c}^{-1}$  (red) are not scattered compared with lower momentum particles (blue and green).  
128 Although there are a certain amount of muons which are scattered on the ground surface  
129 and come from behind the detector, these well-deflected muons are expected to have low  
130 momentum ( $< 10 \text{ GeV c}^{-1}$ ). Besides, this type of upward-going muons were reported by  
131 another dedicated work using scintillator type detectors and TOF analysis (Jourde et al.,  
132 2013). I will discuss these topics furthermore citing this work in the next revision.

133



134

135 **Figure 1 (revised): Deflection angles of CR muons after passing through 100 m**  
 136 **thickness of quartz (a: linear scale, b: logarithmic scale). The injection momenta are**  
 137 **adjusted to the CR muon momentum spectrum for a zenith angle of 80 degrees. The**  
 138 **histogram is divided into three color parts based on the momenta of the ejected muons,**  
 139 **green:  $p_{\text{out}} < 2 \text{ GeV c}^{-1}$ , blue:  $2 \text{ GeV c}^{-1} < p_{\text{out}} < 10 \text{ GeV c}^{-1}$ , and red:  $10 \text{ GeV c}^{-1} < p_{\text{out}}$ .**

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The comment “Unfortunately there are no sufficient experimental data for the angular dependence of the energy spectrum” referring to the electron component, belies the fact that these are electromagnetic residues from calorimetric showers (the atmosphere being the calorimeter) with an energy of 100 MeV. So what do the authors imply ? Do they expect significant variations with energy ? Do they expect significant deviations from the angular distribution of atmospheric muons ? How would this impact their study ?

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While the muon energy spectrum significantly drops due to decay below 1 GeV, the electron spectrum does not drop up to its critical energy in the air (81 MeV). This implies that most of the cosmic electrons at sea level have low energy near 81 MeV and that they have typically large scattering angles in the atmosphere. Thus these well-scattered electrons could hit the detector randomly and are recognized as signals mistakenly. I want to emphasize that these electrons could come from the direction of the mountain, which significantly deteriorates density estimation of the mountain. This is why we think angular distribution of the electron energy spectrum is crucial to muography.

159 In fact, are they making a claim that the ECC detector can remove this kind of events  
160 (“corrupt muon signals”) in a significant and efficient manner ?

161

162 Please see again the revised Figure 1. This shows that the probability that muons deflect  
163 with more than 0.1 rad is extremely small (Figure 1a) and that most of those muons have  
164 momenta less than 2 GeV/c (Figure 1b). Since our ECC module exclude particles with  
165 momenta less than 2 GeV/c, we believe ECC detector exclude most of the corrupt muon  
166 signals.

167

168 Section 3.2 It is true that one can measure on an event by event basis the  
169 momentum of the particle from the measured deflection angles in the detector. This  
170 is really applicable though only for very low energy muons and not low energy  
171 electrons which suffer Bremsstrahlung. Thus a clarifying statement should be made  
172 to how the authors intend to use this information and whether this affects also the  
173 electrons traversing the detector.

174

175 Do you mean by the second sentence that our method is applicable only for very low energy  
176 muons and not applicable for low energy electrons which suffer Bremsstrahlung ? In our  
177 analysis, the energy thresholds were calculated for MIP particles. As you have mentioned,  
178 this is not in the case of electrons. The electrons deflect with larger angles than muons even  
179 if they have the same amount of momentum. But, it does not matter to our purpose since  
180 electrons are noise particles for us. If my answer is not satisfactory, please give us detailed  
181 explanation for your question.

182

183 Section 5.2.3 It would be interesting to know whether any requirement is made for  
184 the ECC detector on possible hits in the first and last layers (i.e. whether a track  
185 must be seen “going out” of the detector or not.

186

187 There is one requirement which is important but was not written in the previous manuscript.  
188 We required tracks in ECC detectors to pass at least five lead plates.

189

190 As for the rest of your comments and language corrections, I will follow your suggestions in  
191 the next revision. Thank you for your careful reading of our manuscript and valuable  
192 questions and comments.

193

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195 **To Referee #2**

196

197 Section 3.2: Multiple scattering is known to be well described by Gaussian for small  
198 deflection angles, but at larger angles it behaves like Rutherford scattering, having  
199 larger tails. The chi-square cut should eliminate tracks in tails. It would be better to  
200 mention it here, or at least to add the word “approximately”.

201

202 I follow your suggestions. In the next revision, I will give more details of deflection angles and  
203 will state that our analysis is based on Gaussian approximation.

204

205 Section 4: The inclination angle  $\theta$  should be related with  $\theta_x$  and  $\theta_y$  by  
206  $\tan(\theta) = \sqrt{\tan(\theta_x)^2 + \tan(\theta_y)^2}$ , not by  $\theta = \sqrt{(\theta_x)^2$   
207  $+ (\theta_y)^2}$  as given in the paper.

208

209 In our analysis, we used the same definition of  $\tan(\theta)$  with your comments. I will change  
210 this part in the next revision.

211

212 Section 5.2.2: "An example of the resultant chi-square distribution is shown in Fig.  
213 6." What is this example? Isn't it data from this experiment? It would be better to  
214 specify what it is or to present the real data from this experiment.

215

216 This resultant chi-square distribution is from our observation of Mt. Showa-Shinzan. I think  
217 the word “example” was misleading. I will change the expression in the next revision.

218

219 Fig. 6: It seems that there are some extra tracks (more than 1 %) in the tail (greater  
220 than the upper bound). Are they so called non-signal tracks? It would be worth  
221 mentioning.

222

223 Due to the error of angular measurement, we have no other way but to regard them as noise  
224 tracks although some of the tracks in the tail region are true muon signals. Thus when  
225 converting the number of particles into the particle flux, we compensate this effect by  
226 multiplying the flux by 100/99. We will mention it in the next revision.

227

228 Section 5.3.1: Are there some tracks which stop inside the ECC detector? If so,  
229 don't they affect the fitting and consequently the efficiency values?

230

231 There should be a certain amount of tracks which stop inside the ECC detector. The possible  
232 candidates for such tracks are low-momentum cosmic electrons (let's say 100 MeV  
233 electrons). However, we could not judge if a particle stops inside our ECC detector or it "looks"  
234 stopped due to inefficiency, because the one-film efficiency of current OPERA film is too low  
235 (~50% as shown in Fig. 8 green). The efficiency fitting was applied to the high-momentum  
236 tracks which survived the severe selection based on scattering angles ( $p > 2 \text{ GeV c}^{-1}$ ). Since  
237 low-momentum cosmic electrons are not selected by our severe selection, the contamination  
238 from electrons in the selected tracks is expected to be low enough to affect the resulting  
239 efficiency values.

240

241 As for the rest of your comments and language corrections, I will follow your suggestions in  
242 the next revision. Thank you for your careful reading and valuable questions and comments.

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### 245 **To Referee #3**

246

247           Nevertheless one could always suggest other sources of background. For instance  
248           there is no mention on the potential effect of ambient radioactivity which is affecting  
249           emulsions.

250

251 We protected emulsion detectors from ambient radioactive particles with the 3-mm-thickness  
252 of steel plates on both sides of the detector (See Tanaka et al. 2007). Figure 9 shows that  
253 ambient beta and gamma rays (< a few MeV) are excluded in our analysis. We will address  
254 this point in the next revision.

255

256           There are other studies already published on the backward upward-going particle  
257           flux which could fake muons coming from the volcano in the absence of time-of-  
258           flight analysis (although it seems that there was a significant amount of rock behind  
259           the emulsions, please precise).

260

261 Jourde et al. (2013) discusses the backward upward-going particles which were found by  
262 scintillation type detectors with a dedicated TOF analysis. This type of particles definitely  
263 cause noise signals also for emulsion detectors. As you have mentioned, there is a significant  
264 amount of rock behind and in front of our emulsion detectors. We think this is why we did not  
265 see an excess of the particle flux in the forward or backward directions. We will address this  
266 issue in the next revision.



267

268

For instance the authors mention a measurement in underground conditions where soft particles barely penetrate and where the same low-value density was obtained with a "Quartet"-type detector. This looks rather suspicious that such low value may be caused by soft particles. More precisions could help clarifying and assessing the main conclusion of the paper which is otherwise of excellent quality.

273

274

I agree with your impression. The noise signals arise not only from soft particles but also from hard particles (muons or pions scattered in the topographic material). I believe the cave measurement will give some constraints on the characteristics of these noise particles. However, details of this work cannot be written here, because it will be published in other place. I will remove this remark on the cave measurement in the next revision, sorry.

279

280

Is it foreseen to couple emulsion detectors with electronic detectors to check instrumental effects with different experimental techniques?

281

282

283

Not yet. But, that sort of study should be conducted near in the future.

284

285

As for the rest of your comments and language corrections, I will follow your suggestions in the next revision. Thank you for your careful reading and valuable questions and comments.

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#### 289 **To Referee #4**

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291

It seems that emulsion film detectors offer critical advantages with respect to detectors based on different principles and the reader is left with the question of why different detectors were employed in the past to accomplish similar tasks (Tanaka et al., 2011; Lesparre et al., 2012).

292

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296

There should be several reasons. For example, it is technically difficult for beginners to construct read-out microscope systems. However, we do not know the exact reason why those precedent studies did not use emulsion films.

297

298

299

300

As discussed in Sect. 5.3.2, the momentum thresholds of these two detectors are: Why do the authors reveal in advance this information, that is discussed later in the paper?

301

302

303

304 This is to show the objectives of our comparative study clearly to readers.

305

306  $\tan_{\theta_y} = 0$  horizontal particles: Are the authors sure that they mean horizontal,  
307 rather than vertical, particles?

308

309 The emulsion detectors were placed perpendicularly to the ground, so the tracks with  
310  $\tan(\theta_y) = 0$  mean horizontal particles.

311

312 As for the rest of your comments and language corrections, I will follow your suggestions in  
313 the next revision. Thank you for your careful reading and valuable questions and comments.

314

315

316 **References (to be added in the next revision):**

317

318 Anastasio et al., The MU-RAY detector for muon radiography of volcanoes, Nuclear  
319 Instruments and Methods in Physics Research A 732 (2013) 423–426.

320

321 Cârloganu et al., Towards a muon radiography of the Puy de Dôme, Geosci. Instrum. Method.  
322 Data Syst., 2, 55-60, 2013.

323

324 Barnafoldi et al. Portable cosmic muon telescope for environmental applications, Nuclear  
325 Instruments and Methods in Physics Research A 689 (2012) 60–69.

326

327 Jourde et al. Effects of upward-going cosmic muons on density radiography of volcanoes,  
328 arXiv:1307.6758.

329