



**Double-sideband
3 mm receiver system**

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A compact receiver system for simultaneous measurements of mesospheric CO and O₃

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Abstract

During the last decades, ground-based microwave radiometry has matured to an established remote sensing technique for measuring vertical profiles of a number of gases in the stratosphere and the mesosphere. Microwave radiometry is the only ground-based technique that can provide vertical profiles of gases in the upper stratosphere and mesosphere both day and night, and even during cloudy conditions. Except for microwave instruments placed at high altitude sites, or at sites with dry atmospheric conditions, only molecules with significant emission lines below 150 GHz, such as CO, H₂O and O₃ can be observed. Vertical profiles of these molecules can give important information about chemistry and dynamics in the middle atmosphere.

Today these measurements are performed at relatively few sites, more simple and reliable instrument solutions are required to make the measurement technique more widely spread. This need is today urgent as the number of satellite sensors observing the middle atmosphere is about to decrease drastically. In this study a compact double-sideband frequency-switched radiometer system for simultaneous observations of mesospheric CO at 115.27 GHz and O₃ at 110.84 GHz is presented

The radiometer, its calibration scheme and observation method are presented. The retrieval procedure, including compensation of the different tropospheric attenuation at the two frequencies, and error characterization are also described. The first measurement series from October 2014 until April 2015 taken at the Onsala Space Observatory, OSO, (57° N, 12° E) is analysed. The retrieved vertical profiles are compared with co-located CO and O₃ data from the MLS instrument on the Aura satellite. The datasets from the instruments agree well to each other. The main differences are the higher OSO volume mixing ratios of O₃ in the upper mesosphere during the winter nights and the higher OSO volume mixing ratios of CO in the mesosphere during the winter. The low bias of mesospheric winter values of CO from MLS compared to ground-based instruments has been reported earlier.

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1 Introduction

Simultaneous measurements of mesospheric gases with different chemical lifetimes, such as ozone (fraction of hour) and carbon monoxide (order of weeks), can give important information on both chemical and dynamical processes in this altitude region.

The middle atmospheric distribution of ozone, O_3 , is characterized by a stratospheric volume mixing ratio (vmr) peak at ~ 35 km altitude, first described by Chapman (1930), and a diurnally varying secondary mesospheric peak at ~ 90 km altitude (Hays and Roble, 1973). The secondary peak is formed during night by reactions between atomic and molecular oxygen and partly destroyed by photo-dissociation during day. Additionally, a tertiary, also diurnally varying, peak is present at ~ 72 km in winter at high latitudes (Marsh et al., 2001; Hartogh et al., 2011).

The main source of middle atmospheric carbon monoxide, CO, is photo-dissociation of carbon dioxide, CO_2 , in the upper mesosphere/thermosphere region. Reactions with hydroxyl, OH, is the main sink. Low vmr in the stratosphere, significantly increasing values with altitude up through the mesosphere, and high values in the thermosphere is the typical vertical distribution of middle atmospheric CO (Lopez-Puertas et al., 2000).

The vertical component of the mesospheric dynamics can at high latitudes be described as an annual cycle with air ascending in the summer and descending in the winter. The horizontal component is weak during summer, while it is controlled by the polar vortex and stronger during winter (Brasseur and Solomon, 2008).

Microwave radiometry is the only ground-based remote-sensing technique that both day and night, even during cloud cover, can provide vertical profiles of different trace gases up to the mesopause region. In microwave radiometry, emission spectra from rotational transitions within the observed molecular species are measured. Due to pressure broadening, the measured spectra contain information about the vertical distribution of the molecule. Except from very dry sites, or sites at high altitudes, only frequencies up to about 150 GHz can be observed since higher frequencies are effectively attenuated by tropospheric water (Janssen, 1993). The gases CO, H_2O , O_2 and

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O₃ all have sufficiently strong emissions at frequencies below 150 GHz. Thus, there is a need for simple and reliable radiometers operating below 150 GHz, since they can observe important gases from almost every ground-based site. If such a radiometer could also observe two of the gases simultaneously (e.g. O₃ and CO) it would be even more useful for the microwave community.

Dicke-switching is the generally used observation technique in microwave radiometry, meaning that the radiation from the sky is compared to an equally intense reference source to diminish the effects of gain variations. Three main Dicke-switching variants can be recognized. In load-switching the reference is a blackbody or other noise source. The zenith sky is the reference in sky-switching. In frequency-switching the mixer's local oscillator frequency, LO, is changed between the signal and the reference phases. Parrish (1994) gives an overview of the mentioned observation methods.

Since the pioneering work by Caton et al. (1968) several heterodyne radiometer systems dedicated for middle atmospheric O₃ observations have been developed, primarily for the relatively strong O₃ transitions at 110.8 and 142.2 GHz. Lobsiger (1987) developed a load-switching technique where the sky, a liquid nitrogen cold load at 80 K, and an ambient load were measured during each observation cycle; several 142.2 GHz instruments use variants of this method (Hartogh et al., 1991; Peter et al., 1998; Studer et al., 2013). Recently the technique has been developed further by implementing a noise diode and a Peltier cooled load (Fernandez et al., 2015).

Nedoluha et al. (2015) use a sky-switching procedure at 110.8 GHz developed by Parrish et al. (1988) and Parrish et al. (1992), where the reference zenith beam passes a “lossy” window at Brewster angle to compensate for the higher intensity in the signal beam.

The advantage of frequency-switching is that the wanted sky emission is present in both signal and reference. The drawback is that the frequency dependent impedances in the frontend components can change the overall gain between the signal and reference phases if the frequency throw is more than ~ 20 MHz. As the pressure broadening in the stratosphere exceeds the bandwidth limitation of frequency-switching this

method can only be used for studies of mesospheric and upper stratospheric O₃. Nagahama et al. (1999) used a frequency throw of 30 MHz and presented vertical O₃ profiles in the altitude range 30–80 km.

Microwave spectra of CO are much narrower than spectra of O₃ due to the different residence altitudes for the two molecules, which make frequency-switching suitable. Waters et al. (1976) made the first microwave CO observations, using absorption measurements against the sun and on-source off-source switching (the standard Dicke method used by radio astronomers). Aellig et al. (1995); Forkman et al. (2003) and Forkman et al. (2012) made frequency-switched observations of CO at 115.3 GHz. Hoffmann et al. (2011) and Straub et al. (2013) used load-switching to observe CO at 230.5 GHz.

The mixer is the key component in the heterodyne radiometer. The incoming radio frequency, RF, is mixed with the LO, and the output intermediate frequency, IF, is a mix of the upper and lower sidebands. To avoid the unwanted sideband (or image band) the radiometer can be operated in single sideband mode where the image band is suppressed before the mixing. If none of the sidebands are suppressed, we have a double-sideband system which makes it possible to observe signals from the two bands simultaneously. The disadvantages are that the sideband ratio has to be known and the tropospheric attenuation has to be corrected individually for the two bands if the tropospheric opacity differ between the two frequencies. Except for instruments where the LO is placed on the center of the observed line, e.g. the 183 GHz water vapor radiometer for the ALMA project (Emrich et al., 2009), most ground-based radiometers today are single sideband instruments. One exception is the 110–116 GHz radiometer for CO and O₃ observations designed and operated by Pidtyachiy et al. (2010).

In this study we present the first simultaneous measurements of mesospheric O₃ at 110.8 GHz and CO at 115.3 GHz made by a ground-based, double sideband and frequency-switched radiometer system. The system is operated at the Onsala Space Observatory, OSO, (57.4° N, 11.9° E). The instrument, its calibration scheme, the retrieval procedure and the first results are introduced. Section 2 describes the receiver

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system and the calibration and Sect. 3 presents the inversions. The results are given in Sect. 4 and the error analysis is described in Sect. 5. Section 6 shows a satellite comparison, and Sect. 7 gives a summary and the conclusions.

2 Instrument and observation technique

We present a double-sideband, frequency-switched heterodyne receiver system for simultaneous spectral measurements of the atmospheric O_3 $6_{15} \rightarrow 6_{06}$ transition at 110.83604 GHz and the CO $1 \rightarrow 0$ transition at 115.27120 GHz, observed at an elevation of 80° . A block diagram of the receiver is shown in Fig. 1 and technical specifications are given in Table 1. A 2 bit autocorrelator is used as backend spectrometer. The bandwidth is 20 MHz and the nominal resolution is 25 kHz (800 delay channels).

2.1 Frontend description

The receiver frontend includes a four stage Low Noise Amplifier, LNA, a fundamental resistive mixer, and a $\times 4$ LO chain, all integrated onto a single Monolithic Microwave Integrated Circuit, MMIC, using a 100 nm mHEMT process. The mixer provides a conversion loss of 8–10 dB for LO power of 4 dBm. The LO chain consists of two doublers followed by a two stage power amplifier. The amplifier delivers about 5 dBm of LO signal to the mixer with an input power of 9 dBm at 29.5 GHz. Vassilev et al. (2010) gives more details on the performance of the receiver and a breakout of the LNA.

2.2 Calibration

Brightness temperature, T_b , derived from the Rayleigh–Jeans approximation of the Planck law, is often used as a measure of the received radiation in microwave radiometry. The Rayleigh–Jeans approximation can be written:

$$B(\lambda, T) \approx \frac{2kT}{\lambda^2} \quad (1)$$

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where B is the brightness describing the energy emitted by a black body, λ the wavelength, k the Boltzmann constant, and T the physical temperature of the black body. Equation (1) is valid when $h\nu \ll kT$, where h is the Planck constant and ν is the frequency. The brightness temperature, T_b , is defined as:

$$T_b = I(\lambda) \frac{\lambda^2}{2k} \quad (2)$$

The proportionality between the received radiation, I , and T_b is the reason why T_b is used in microwave radiometry. The antenna temperature, T_a , is defined as the convolution between the observed brightness temperature distribution and the antenna pattern. In the rest of this section a pencil beam is assumed implying that the measured antenna temperature, T_a , is equal to the brightness temperature, T_b , in the observed direction.

The system temperature, the radiometer output power measured by the spectrometer, is defined as $T_{\text{sys}} = T_a + T_{\text{rec}}$, where the receiver temperature, T_{rec} , is a measure of the power generated in the components along the radiometer system transmission line where the first stages as LNA and mixer contribute the most.

In the mixer, the RF input spectrum is folded around the LO to form the IF output spectrum (see Fig. 2). The IF bandpass filter selects the position and width of both the lower sideband, LSB, and the upper sideband, USB. If any of the two sidebands are terminated ahead the mixer the receiver is called single sideband, SSB. We use the mixer in true double sideband mode, DSB, where LSB is centered at 110.84 GHz and USB at 115.27 GHz, see simulated spectra in Fig. 3. The contributions from LSB and USB are weighted with their relative frontend gains and then added to form T_{sys} (Ulich and Haas, 1976). The system temperature of a calibration blackbody load that fills the antenna beam, $T_{\text{sys}}(\text{load})$, can thus be expressed as:

$$T_{\text{sys}}(\text{load}) = G_L (T_{\text{load}}(L) + T_{\text{rec}}(L)) + G_U (T_{\text{load}}(U) + T_{\text{rec}}(U)) \quad (3)$$

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where L and U mark the contributions from the LSB and USB frequencies, G_L and G_U are the normalized relative frontend power gains ($G_L + G_U = 1$) in the two sidebands (also called sideband responses), and T_{load} is the temperature of the blackbody load. The sum of the two contributions to the receiver temperature is denoted T_{rec} , i.e. $T_{\text{rec}} = G_L T_{\text{rec}}(L) + G_U T_{\text{rec}}(U)$. If it is assumed that the load is a blackbody in both sidebands, Eq. (3) is hence simplified to:

$$T_{\text{sys}}(\text{load}) = T_{\text{load}} + T_{\text{rec}} \quad (4)$$

To estimate T_{rec} two blackbody loads with physical temperatures T_{hot} (ambient load) and T_{cold} (77 K load) are observed each month. T_{rec} can then be estimated using:

$$\frac{P_{\text{cold}}}{P_{\text{hot}} - P_{\text{cold}}} = \frac{T_{\text{sys}}(\text{cold})}{T_{\text{sys}}(\text{hot}) - T_{\text{sys}}(\text{cold})} = \frac{T_{\text{cold}} + T_{\text{rec}}}{T_{\text{hot}} - T_{\text{cold}}} \rightarrow T_{\text{rec}} = P_{\text{cold}} \frac{T_{\text{hot}} - T_{\text{cold}}}{P_{\text{hot}} - P_{\text{cold}}} - T_{\text{cold}}, \quad (5)$$

where P_{hot} and P_{cold} are the measured powers observing the two loads. The system temperature when observing the sky, $T_{\text{sys}}(\text{sky})$, is given by:

$$T_{\text{sys}}(\text{sky}) = G_L T_a(L) + G_U T_a(U) + T_{\text{rec}} \quad (6)$$

where $T_a(L)$ and $T_a(U)$ are the antenna temperatures at 110.84 and 115.27 GHz, respectively.

The following calibration procedure is performed each 15 min to estimate the sky brightness temperature:

$$\frac{P_{\text{load}} - P_{\text{sky}}}{P_{\text{sky}}} = \frac{T_{\text{sys}}(\text{load}) - T_{\text{sys}}(\text{sky})}{T_{\text{sys}}(\text{sky})} = \frac{(T_{\text{load}} + T_{\text{rec}}) - (G_L T_a(L) + G_U T_a(U) + T_{\text{rec}})}{G_L T_a(L) + G_U T_a(U) + T_{\text{rec}}}, \quad (7)$$

where P_{load} and P_{sky} are the measured powers observing the load and the sky, respectively. The weighted mean of the antenna temperatures at the two sidebands,

$$T_{\text{sky}} = G_L T_a(L) + G_U T_a(U) \quad (8)$$

can be derived from Eq. (7) since T_{load} and T_{rec} are known. Since a pencil beam is assumed, T_{sky} is denoted as sky brightness temperature (see above). An error in the estimate of T_{rec} introduces an error in the estimation of T_{sky} . The hot-cold calibrations (Eq. 5) performed so far indicate that the variation in T_{rec} is less than 3%. Equations (7) and (8) then gives that the error in T_{sky} is less than 2%.

The sky brightness temperature at 115.3 GHz is 35–60 K higher than at 110.8 GHz. This is explained both by the frequency variation of absorption due to tropospheric water and by the fact that 115.3 GHz is situated higher on the wing of the 118 GHz O_2 line, see the broadband spectra in Fig. 4 estimated from one year of radiosonde data taken at Landvetter Airport, 38 km N.E. of Onsala Space Observatory.

2.3 Frequency-switching

The particular Dicke-switch method used here is frequency-switching. In this method the frontend mixer LO frequency, f_{LO} , is switched between the phases of the signal, S , and the reference, R , in the measurement cycle. $f_{\text{LO}}(S) = f_c - \Delta f$ and $f_{\text{LO}}(R) = f_c + \Delta f$ where f_c is the mean of the two local oscillator frequencies. Owing to $S - R$ being a difference, the spectra will show both a negative and a positive peak in the observed spectral characteristic, with a separation equal to the frequency throw, $2\Delta f$. An averaged spectrum is seen in Fig. 5. The spectrum is a combination of double-sideband measurement and frequency-switching which explains the positions of the negative and positive peaks of O_3 from the lower and CO from the upper sidebands.

Using frequency-switching during the observation cycle we record

$$\begin{aligned} \Delta T_{\text{sky}} &= \frac{\Delta P}{P_{\text{load}} - P_{\text{sky}}} (T_{\text{load}} - T_{\text{sky}}) = T_{\text{sky}}(-\Delta f) - T_{\text{sky}}(+\Delta f) \\ &= G_L (T_a(L, -\Delta f) - T_a(L, +\Delta f)) + G_U (T_a(U, -\Delta f) - T_a(U, +\Delta f)) \end{aligned} \quad (9)$$

where ΔT_{sky} is the difference in brightness temperatures (since we assume a pencil beam) and $\Delta P = P_S - P_R$ is the difference in the measured powers between the two

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frequencies $f_{LO}(S)$ and $f_{LO}(R)$. The calibration procedure gives P_{load} , P_{sky} , T_{load} and T_{sky} .

2.4 Sideband ratio

The sideband responses G_L and G_U have to be known accurately to be able to retrieve vertical profiles from the measured spectra. The measurement of the sideband responses relies on inserting a continuous wave (CW) of known amplitude in the RF path of the instrument and to measure the down converted IF signal. The measurement is then repeated over several RF/IF frequencies to get the overall sideband response.

Figure 6 shows the setup used for the measurement of the sideband response of the instrument. A millimeter wave source generates a CW in the 110.5–115.5 GHz frequency band. A mm-wave spectrum analyzer extender measures the amplitude of the CW signal. The radiometer frontend, and a spectrum analyzer measure the amplitude of the down converted IF. All the measurement equipments are synchronized to a common reference clock.

In the current measurement setup, the mm-wave source is first connected to the extenders to measure the amplitude of the CW signal while sweeping the source frequency $f_{RF} = 110.5 \rightarrow 115.5$ GHz. After taking the RF power sweep, the CW source is connected to radiometer, and the amplitude of the down converted IF is recorded while sweeping the RF frequency as before. The local oscillator frequency of the radiometer is held constant at $f_{LO} = 113.055$ GHz (28.26375×4). The two sidebands after the IF amplifier bandpass response are,

$$\begin{aligned} f_{LSB} &= 113.055 - (1.5 \rightarrow 2.5) = 110.555 \rightarrow 111.555 \text{ GHz} \\ f_{USB} &= 113.055 + (1.5 \rightarrow 2.5) = 114.555 \rightarrow 115.555 \text{ GHz} \end{aligned} \quad (10)$$

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The sideband gains of the instrument can then be estimated by taking the ratios of the measured power at RF frequencies and IF frequencies as,

$$g_{\text{LSB}} = \frac{P_{\text{IF}}^{\text{LSB}}}{P_{\text{RF}}^{\text{LSB}}} \quad \text{and} \quad g_{\text{USB}} = \frac{P_{\text{IF}}^{\text{USB}}}{P_{\text{RF}}^{\text{USB}}} \quad (11)$$

The measured sideband ratio, $g_{\text{LSB}}/g_{\text{USB}}$, is close to unity. However, standing-wave patterns are seen in both the measured RF and IF powers, which introduce an error in our estimation of the sideband ratio. These standing waves have to be minimized in order to improve the quality of the measurements. Nevertheless, the results obtained so far with the current setup are promising. The linear normalized relative frontend power gains, G_L and G_U are estimated at 0.5 ± 0.05 and 0.5 ∓ 0.05 respectively.

2.5 Water vapor radiometer for tropospheric measurements

Due to the nature of DSB mixers, the measured sky brightness temperature is the mean of the brightness temperatures at the lower and upper sideband frequencies, weighted with their respectively sideband gains G_L and G_U . To be able to correct for the tropospheric attenuation an estimation of the sky brightness temperatures at these two frequency regimes are needed. The OSO site operates two dual-frequency radiometers, ASTRID and KONRAD, that continuously measure the sky brightness temperature in different directions at 21.0/31.4 and 20.6/31.6 GHz respectively, see Table 2. The data is used to provide independent corrections for the water vapor induced time delay which affect the accuracy of the geodetic VLBI observations performed at the observatory (Elgered and Jarlemark, 1998). In Sect. 3.3 the use of the calibrated zenith sky brightness temperatures, from these instruments, to estimate the tropospheric opacity at 110.84 and 115.27 GHz, is discussed.

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3 Retrievals

3.1 Forward model

For the retrievals presented in this paper, the Atmospheric Radiative Transfer Simulator (ARTS v.2.3.145) is used as a forward model (Buehler et al., 2005; Eriksson et al., 2011). It is a general radiative transfer model that can provide Jacobians for a large number of different measurement geometries and systems. A 1-D simulation setup is applied using a pressure grid ranging from 1.3×10^5 Pa (0 m) to 7.5×10^{-4} Pa (~ 130 km) with a spacing of ~ 250 m. Line-by-line simulations of frequencies in two bands between 110.816–110.856 and 115.251–115.291 GHz are run with a monochromatic frequency grid having a spacing of 4.2 MHz at the far end of each band, decreasing to 14.13 kHz in the center of each band. The instrument is modeled as a dual sideband receiver with a flat 50 % sideband response in each band. Each channel of the autocorrelator is modeled to have a channels response corresponding to an ideal Hanning filter with a FWHM of 50 kHz. The antenna is modeled as a pencil beam antenna looking at a zenith angle of 10° , and the instrument is positioned at ground level.

The spectroscopic lines included in the forward model are CO at 115.27 GHz, O₃ at 110.5, 110.7, 110.8 and 115 GHz as well as complete absorption models for molecular oxygen, molecular nitrogen, water vapor and liquid water. A summary of the spectroscopic parameters is given in Tables 3 and 4. For the emission lines the line position and strength is taken from JPL-catalog (Pickett et al., 1998) while the broadening parameters are taken from HITRAN04 (Rothman et al., 2005).

Comparing the measurements to a forward model simulation with the line positions from the JPL-catalogue (as well as HITRAN04 Rothman et al., 2005) the simulated CO emission occurs at the same frequency in both the simulation and our measurements, while the simulated O₃ line emission at 110.8360400 GHz shows a clear frequency offset compared to the measurements. Since the CO line is positioned correctly a shift in the LO frequency cannot explain the frequency offset of the O₃ line. This thus indicates that the databases have the wrong frequency for this spectral line. Best agreement be-

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tween the forward model and measurement was found if the line was shifted 117 kHz (specified uncertainty is 50 kHz) to 110.8359230 GHz. Note that for the purpose of this study, the exact reason for this shift is not relevant, since a pure shift in frequency does not affect the retrieved concentrations as long as the modeled and measured spectra are consistent.

3.2 Retrieval model

To retrieve CO and O₃ concentrations from the measured spectra, the maximum a posteriori method, also called optimal estimation method, OEM, (Rodgers, 2000) is used as implemented in the updated version of the Qpack software (Eriksson et al., 2005).

Given the spectra with assumed errors and a statistical distribution of the measured atmosphere, the method returns the maximum a posteriori estimate combining these two pieces of information. If the atmosphere and possible instrument parameters are described by a state vector \mathbf{x} , the measured spectrum by \mathbf{y} , and the a priori atmosphere as \mathbf{x}_a , the estimated atmosphere is

$$\hat{\mathbf{x}} = \mathbf{x}_a + (\mathbf{K}^T \mathbf{S}_e^{-1} \mathbf{K} + \mathbf{S}_a^{-1})^{-1} \mathbf{K}^T \mathbf{S}_e^{-1} (\mathbf{y} - \mathbf{K} \mathbf{x}_a), \quad (12)$$

where \mathbf{S}_e , \mathbf{S}_a are the covariance matrices describing the uncertainty (assuming normal distribution) in the measurements and a priori atmosphere respectively. The Jacobian or Weighting function matrix, $\mathbf{K} \equiv \partial \mathbf{y} / \partial \mathbf{x}$, is the linearized derivative of the forward model and describes how a change in any of the state vector elements influences the measured spectrum.

To save computational resources, the inverse problem (Eq. 12) is solved on a coarser grid than the forward model. The state vector is specified to contain the concentration (vmr) of CO and O₃ at pressure levels between 1×10^5 and 1×10^{-3} Pa with a spacing of 2 km. In addition to CO and O₃ the state vector includes the concentration (in units relative to the a priori concentration) of water vapor and liquid water between 1×10^5 and 1.3×10^3 Pa with a spacing of 1 km. These species are included to correct for tropospheric influence on the mesospheric emission (see Sect. 3.3). The elements of the

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state vector containing these species are referred to as \mathbf{x}^{trop} . To account for baseline ripple in the instrument a 3rd order polynomial fit is performed, and its coefficients are stored in the four last elements of the state vector.

Each of these state vectors variables needs a priori values stored in \mathbf{x}_a . The a priori profile for CO and O₃ is based on a climatology containing the monthly zonal mean values from ACE-FTS at 57.5° N. It is based on the method described in Jones et al. (2012) but with an updated data quality classification (Sheese et al., 2015). The climatology covers pressure levels from 1×10^5 to 1×10^{-4} Pa, but lacks data for certain months and altitudes. A linear interpolation between months is used if values are missing. Above 1×10^{-4} Pa the climatology is extrapolated using the vmr value from 1×10^{-4} Pa. The temperature, altitude and pressure relationship is, above 100 Pa, taken from a climatology based on the MSISE-90 model (Hedin, 1991), while below 5000 Pa it is based on the database for used tropospheric correction (see Sect. 3.3). Between 5000 and 100 Pa the temperatures are obtained by a linear interpolation between the two datasets.

To solve Eq. (12), \mathbf{S}_e and \mathbf{S}_a must be specified. We describe these covariances with a standard deviation and a correlation function (see e.g. Christensen and Eriksson, 2013). For \mathbf{S}_e the standard deviation is equal to the thermal noise estimated from the measurements (~ 0.07 K) and correlation between channels is modeled as a Gaussian correlation function with a correlation length equal to 1.6 channels. The specification of \mathbf{S}_a depends on which state vector variable the elements describe. The covariance of CO is described with a standard deviation equal to 1 ppmv plus 100 % of the annual mean CO profile (ppmv) from the a priori. This large uncertainty is needed to ensure a reasonable sensitivity despite the low signal to noise ratio of the CO measurements. For O₃ the standard deviation is described simply as 4 ppmv for all altitudes, and for temperature it is set to 5 K for all altitudes. The correlation between altitudes is set to follow a linear correlation function with a correlation length of 8 km for both species and the temperature. Tropospheric water vapor has a standard deviation equal to 10 % of the a priori value and a linear correlation function with a correlation length of 8 km,

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while the liquid water has a standard deviation equal to 100 % of the a priori value, and no correlation between altitudes. The baseline fit has a standard deviation of 4 K for all coefficients.

3.3 Tropospheric correction

In order to accurately estimate the CO and O₃ concentrations in the mesosphere, the tropospheric attenuation needs to be accounted for. A common way of to achieve this is to model the troposphere as a single layer, with an effective temperature and opacity, and performing a correction of the observed spectra prior to performing the retrieval. For the DSB receiver the difference in the opacity between the two sidebands is too large for such an approach to work, and as such the troposphere needs to be included directly into the forward model. This is done in two steps. First an atmosphere is selected from a database of tropospheric scenarios. The atmosphere selected is the one minimizing the following cost function

$$\chi^2 = (\mathbf{y}^{\text{trop}} - f(\mathbf{x}^{\text{trop}}))^T \mathbf{S}_e^{\text{trop}} (\mathbf{y}^{\text{trop}} - f(\mathbf{x}^{\text{trop}})), \quad (13)$$

where \mathbf{y}^{trop} is the measurements used for the tropospheric correction, $f(\mathbf{x}^{\text{trop}})$ the radiance from the modeled troposphere and $\mathbf{S}_e^{\text{trop}}$ the covariance matrix describing the measurement noise for the measurements used for the tropospheric retrieval. For the DSB instrument \mathbf{y}^{trop} consists of two elements, the mean T_{sky} measured across all channels, $T_{\text{sky}}^{\text{mean}}$, and the ground temperature at OSO at the time of the measurement, T_{ground} , measured by the weather station at the site. Both these values are averaged over the same time period as the spectral measurements. The second step is to expand \mathbf{y} in Eq. (12) to include \mathbf{y}^{trop} and retrieve \mathbf{x}^{trop} with the OEM method, using the selected troposphere as the a priori. The effect of tropospheric attenuation on the mesospheric spectra are thus also added to \mathbf{K} .

The tropospheric states considered in Eq. (13) are taken from a database constructed on data from the ERA-Interim project (Dee et al., 2011), covering years 2009,

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2010 and 2011. Data were extracted for the OSO site, at 00:00 and 12:00 UTH each day of the 3 year long period. In total, the database contains 2190 atmospheric states. Temperature and humidity data were used as provided by ERA-Interim. In a first step, LWC was set to zero if $r > f$, where f is the ERA-interim cloud fraction and r a random number with flat probability between 0 and 1. If $r \leq f$, LWC was set to LWC_0/f , where LWC_0 is the ERA-Interim grid box mean LWC. However, in comparison with ASTRID and KONRAD retrievals, this was found to underestimate both the amount of cloudy situations and maximum integrated liquid water (LWP). A better agreement was created artificially by instead selecting r between 0 and 0.5, and increasing LWC_0 with a random factor, uniformly distributed between 1 and 2.

Fitting the troposphere using just T_{sky}^{mean} and T_{ground} is a grossly under-determined problem, and thus to test the accuracy of this method the tropospheric attenuation was also simultaneously estimated by including measurements from the water vapor radiometer ASTRID into y^{trop} (averaged over the same timeperiod as the double sideband receiver). The two channels are simulated as described in Sect. 2.5 using the same settings as described in Sect. 3.1. Simulations were also run for KONRAD, and comparing the simulated brightness temperatures from the two water vapor radiometers and brightness temperature measured, an offset was seen. For clear sky days (i.e. no clouds) ASTRID systematically measured brightness temperatures 3 and 5 K lower than the simulations predicted for the lower and upper frequency channels respectively, while KONRAD had a bias of -2 and $+1$ K for the two channels. Since both radiometers differ in their bias, we assume that this discrepancy comes from instrumental errors. For this study, the ASTRID instrument alone is used to characterize possible errors in the tropospheric correction described by Eq. (13) (see Sect. 5). Thus, in order to ensure consistency between the simulations and the measurements, the ASTRID was bias corrected to match the simulated data before it is used.

4 Results of the OSO measurements

Figure 7 shows retrieved profiles from two example cases, one captured during a winter night and one during a autumn day. The winter spectrum shows stronger emission and less noise than the summer spectrum. The different noise levels mainly come from the higher tropospheric humidity in the autumn leading to more attenuation of the mesospheric signal. Figure 8 shows all retrieved CO and O₃ from the measurement period. From the retrieved profiles, it is clear that stronger emission in the winter come from an increase in CO and O₃ at altitudes above 10 Pa. The general structure of the CO distribution is seen in Fig. 8 with a sharp increase in volume mixing in the upper mesosphere.

This initial time series mainly cover the winter period. During the winter the general circulation brings down air from the thermosphere into the mesosphere which increases the mesospheric CO abundance. This down-welling is strongest inside the polar vortex, and the variation of CO seen from day to day is mainly explained by movement of the polar vortex. OSO is sometimes located within and sometimes outside the vortex during the winter. During summer the general circulation is reversed and hence is characterized by up-welling air which decrease the mesospheric CO abundance, which can be seen in the end of the measurement series.

The time series of O₃ both show the upper part of the stratospheric peak and a nighttime peak at altitudes above 10 Pa during the winter. Due to the poor resolution of the instrument the observed mesospheric diurnal peak can be a mixture of both “the secondary ozone peak” at ~ 90 km and “the tertiary ozone peak”, located at 72 km, see Sect. 1.

Example averaging kernels are shown in Fig. 9. The large negative values for the higher altitude CO averaging kernels with respect to changes in vmr at lower altitudes are common in ground based microwave radiometers retrieving CO (Hoffmann et al., 2011; Forkman et al., 2012; Straub et al., 2013). However, since the probability of a large change in volume mixing ratio at lower altitudes is small, these negative values

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are not problematic. The retrievals have a measurement response above 0.8 between 100 and 0.2 Pa for CO and 200 to 0.6 Pa for O₃. Calculating the degrees of freedom of the retrievals (trace of averaging kernel matrix) CO is retrieved with 1.5–2.5° of freedom (depending on season) and O₃ with 3.5–5° of freedom. Resulting to an average vertical resolution of 20 and 10 km for CO and O₃ respectively. For the highest altitudes (e.g. 0.1 Pa), the peak of the averaging kernel also deviates from the altitude which the averaging kernel represents.

5 Sensitivity to errors in forward model and retrieval parameters

Errors are introduced from uncertainties in the forward model and the retrieval parameters. These include uncertainties in the modeling of the instrument, uncertainties in the spectroscopic parameters used, uncertainties in the tropospheric correction as well as a dependence on the a priori assumptions used in the retrievals. To estimate these errors the retrievals are rerun with each parameter perturbed with its 1σ uncertainty (see Table 5). For the tropospheric correction the error was estimated by comparing the nominal correction method (using only $T_{\text{sky}}^{\text{mean}}$ and T_{ground}) to the extended tropospheric correction including ASTRID. The error estimation was carried out over the sub-set of measurements where simultaneous data from the OSO instrument and ASTRID was available.

The mean difference and the standard deviation between each of the perturbed retrievals and the standard retrievals are shown in Fig. 10, together with the total root-sum-square error from all the parameters. For CO the estimated accuracy is better than 0.3 ppmv for altitudes below 10 Pa, while degrading at altitudes above this, to 2 ppmv at 1 Pa. The largest source of systematic uncertainty is the characterization of the sideband response, followed by uncertainties in the temperature profile. The total estimated random errors for CO from the retrieval parameters are of the same size as the random error from thermal noise in the measurements (< 0.3 ppmv below 10 Pa and 1 ppmv at 1 Pa). For O₃, the estimated systematic and random uncertainties from the simulated

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error sources, are better than 0.5 ppmv between 100 and 1 Pa, with the largest source of systematic errors being the uncertainties in the sideband response and the a priori variability. Errors due to thermal noise in the measurements are 0.5 ppmv across all altitudes where the measurement response is greater than 0.8. A summary estimated precision and accuracy for the retrieved data at three example pressure levels is given in Table 6. For the precision estimate, the maximum error from either thermal noise or forward model parameters is used.

6 Satellite comparisons

The vertical profiles from the OSO instruments have been compared to version V-3-3 of CO and O₃ data from the microwave limb sounder MLS on the Aura satellite (Pumphrey et al., 2007) and (Froidevaux et al., 2008), see Table 7. The comparison covers the time period October 2014 until April 2015. MLS data taken closer to the OSO-site than latitude ±5° and longitude ±10° has been used, see Fig. 11. MLS has measurements solely from either night (01:00–02:00 UTC) or day (11:00–12:00 UTC) within the used position range. Since the OSO data is 6 h averages the 6 h period with the best overlap with the MLS measurement times has been used in the comparison. The MLS data was interpolated onto the OSO retrieval grid. To compensate for the different vertical resolution of the two instruments the MLS data was convolved with the averaging kernels, A, of the OSO instrument (Rodgers and Connor, 2003)

$$\mathbf{x}_s = \mathbf{x}_a + \mathbf{A}(\mathbf{x}_{\text{sat}} - \mathbf{x}_a), \quad (14)$$

where \mathbf{x}_a is the OSO a priori profile and \mathbf{x}_s is the smoothed MLS profile.

Figures 12 and 14 show mean profiles of O₃ and CO for the two instruments. Figure 12 shows averaged night and day O₃ profiles from December 2014 and Fig. 13 shows the difference in vmr between OSO and MLS. The averaged day profiles from the two instrument are very similar within their measurement ranges. The night profiles however differ at altitudes above 5 Pa (~70 km), where OSO shows a more pro-

nounced peak in the upper mesosphere. The MLS peak seen in the night profile at 2 Pa is probably “The tertiary ozone peak”.

There is no clear diurnal variation of the CO profiles. Figure 14 shows averaged day profiles from December 2014 and March 2015 and Fig. 15 shows the difference in vmr between OSO and MLS. OSO shows higher CO abundances than MLS at altitudes above 5 Pa during December. During March the difference between the two instruments is much less pronounced.

Figures 16 and 17 show time series for the measurement period for OSO and MLS at three different pressure levels (100, 10, and 1 Pa). The average measurement response for OSO is higher than 80 % for both O₃ and CO at these pressure levels and MLS reports valid data at altitudes below ~ 1 Pa. Note that due to the vertical resolution of the OSO instrument, the values at these pressure levels are not necessarily completely independent. The CO and O₃ data from the two instruments shows the same general features, both in terms of the overall variation and in sporadic events. The main differences between the two instruments are both the higher OSO values of upper mesospheric O₃ mixing ratios during winter nights and the higher OSO values of upper mesospheric CO mixing ratios during the winter compared to MLS (see also Figs. 12 and 14).

Similar discrepancies between CO measurements from MLS and ground-based instruments, as presented above, have been reported earlier by Forkman et al. (2012) using an older receiver system and by Hoffmann et al. (2011).

7 Summary and conclusions

The first simultaneous measurements of mesospheric O₃ at 110.8 GHz and CO at 115.3 GHz made by a ground-based, double sideband and frequency-switched radiometer system operated at the Onsala Space Observatory, OSO, (57.4° N, 11.9° E) are presented.

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Dicke-switching is the generally used observation method in microwave radiometry to diminish effects of gain variations in the receiver system. Frequency-switching is the most time effective Dicke-switching variant since no reference load is observed except in the calibrations. Since the frequency throw has to be less than ~ 20 MHz to avoid gain differences, the method is restricted for studies of the spectral shapes of emission lines from high altitudes where the pressure broadening is limited. The method is hence well-adapted for observations of mesospheric CO and O₃.

Most ground-based microwave heterodyne radiometers for atmospheric remote sensing are operated in single sideband mode. In a double sideband system simultaneous measurements of two emission lines at rather different frequencies, as O₃ at 110.84 GHz and CO at 115.27 GHz, are possible. The drawbacks of a system where both sidebands are used are both that the sideband ratio has to be measured and that the tropospheric attenuation can differ between the two line frequencies.

In this study the gain between the frontend RF input and IF output was estimated by measuring the IF power when a calibrated RF source was connected to the frontend. The RF source was swept across the lower and upper sidebands and the sideband ratio was estimated by comparing the IF and RF powers in the measured frequency range. Standing waves arising from reflections in the transmission line affects the result. In order to reduce the reported error in the sideband ratio estimation, the measurement setup will be refined to try diminish the standing waves.

The commonly used method to compensate measured spectra for the tropospheric attenuation is to use an one-layer model of the troposphere with constant effective temperature and opacity and to correct the observed spectra before the retrieval process. The difference between the opacities in the two sidebands is however too large for this method to work. A new approach where the troposphere is included in the forward model has been developed.

To calculate vertical profiles of CO and O₃ from the measured spectra the Optimal Estimation Method, OEM, has been used in the retrieval process. To present as exact

error estimations as possible, the systematic effects arising from the uncertainties in the different measurement and retrieval parameters, have been carefully studied.

The OSO CO and O₃ data have been compared to measurements from the satellite instrument MLS (v3-3) on Aura. The data from two instruments shows the same general features in both sporadic events and in the overall variation. The main differences between the instruments are the higher OSO values of O₃ mixing ratios in the upper mesosphere during the winter nights and the higher OSO winter values of CO mixing ratios in the upper mesosphere compared to MLS.

Microwave radiometry is the only ground-based remote sensing technique that can monitor the mesosphere day and night even during cloudy conditions. Simple and reliable microwave radiometers measuring in the frequency range below 150 GHz can be very valuable for mesospheric research since they can be operated at almost every ground-based site. The described instrument shows the potential of a double-sideband and frequency-switched radiometer system for simultaneous measurements of mesospheric CO and O₃.

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Radio frequency	110–116 GHz, DSB
Mirror edge taper	−35 dB
Elevation	80°, fixed
Horn	Aluminium, corrugated
Beam width, FWHM	6°
First stage	LNA +20 dB, Ambient temperature
Image sideband rejection	None, DSB
Sideband response	0.50/0.50 ± 0.05
Local oscillator (LO)	Synth. + multipliers
LO frequency	113 GHz
Frequency throw (2Δ <i>f</i>)	8 MHz
Mixer IF	2.21 GHz
DSB receiver temperature	~ 450 K
Backend spectrometer	800 channel autocorrelator
Bandwidth	20 MHz
Nominal resolution	25 kHz
Integration time	6 h centered at 05:00, 11:00, 17:00, and 23:00 UTC

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Table 2. Specifications for the total power dual channel radiometers.

Radiometer	ASTRID	KONRAD	Unit
Radio frequencies	21.0/31.4	20.6/31.6	[GHz]
Antenna (one for each frequency)	Dielectrically loaded horn	Conical lens horns	
Beam width, FWHM	6/6	2.9/2.0	[°]
Pointing resolution	0.1	0.1	[°]
Reference load temperatures	313/360	313/373	[K]
System noise temperature	450/550	450/550	[K]
RF bandwidth (both channels)	1000	320	[MHz]
Accuracy	< 1	0.5	[K]

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Table 3. Summary of the two major spectroscopic lines.

Line parameter	CO	O ₃	Unit
Center frequency, f_0	115.2712018	110.8359230	[GHz]
Line intensity, I_0	9.761128×10^{-18}	3.567796×10^{-17}	[m ² Hz ⁻¹]
Air broadened width	23 332.68	23 932.87	[Hz Pa ⁻¹]
Self broadened width	25 958.54	30 009.87	[Hz Pa ⁻¹]
Ref. temp. for broad. param.	296	296	[K]
Temp. dep. exp. for broad. param.	0.69	0.73	[-]

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Table 4. Summary of the complete absorption models. The model name refers to the name used internally in ARTS, while the model is described in the reference given.

Species	Absorption model	Reference
N ₂	N2-SelfContStandardType	Rosenkranz (1993)
O ₂	O2-PWR98	Rosenkranz (1998)
H ₂ O	H2O-PWR98	Rosenkranz (1998)
LiquidWater	liquidcloud-MPM93	Liebe et al. (1993)

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Table 5. Summary of the perturbations applied to the forward model and retrieval parameters in the sensitivity study. Method indicates how the perturbation values were estimated.

Parameter	Perturbation (1σ)	Method
Line strength (O_3/CO)	2%	Pickett et al. (1998)
Pressure broadening parameter (O_3/CO)	10%	Rothman et al. (2005)
A priori profile	50%	–
A priori uncertainty	50%	–
Temperature profile	± 5 K	1σ of MSISE-90 is 3 K (Hedin, 1991)
Sideband response	5%	Sect. 2.4
Tropospheric correction	comparison to method	using ASTRID

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Table 6. Summary of error estimate.

Species	Error (ppmv)	100 Pa	10 Pa	1 Pa
O ₃	Precision	0.31	0.44	0.49
O ₃	Accuracy	0.68	0.15	0.28
CO	Precision	0.19	0.29	1.30
CO	Accuracy	0.07	0.24	1.78

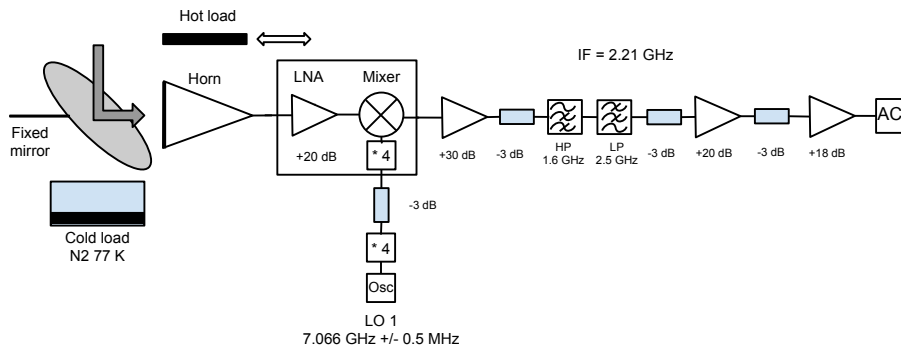


Figure 1. Block diagram of the DSB 110–115 GHz O₃/CO receiver system. The cold load is regularly mounted and used.

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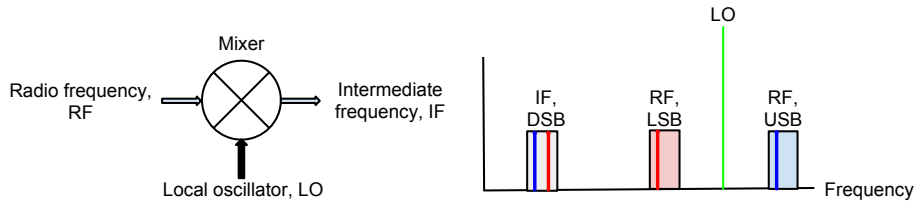


Figure 2. Mixer fundamentals.

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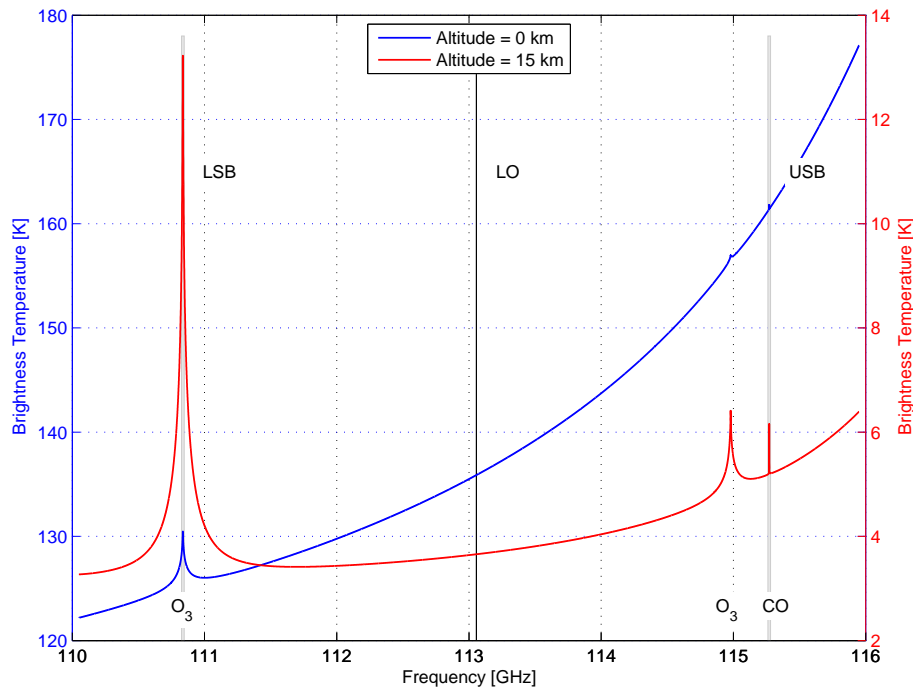


Figure 3. Simulated atmospheric spectra as seen with an elevation of 80° from the ground (blue) and, for clarity, from an altitude of 15 km (red). The 20 MHz wide LSB and USB frequency ranges and the LO frequency are marked.

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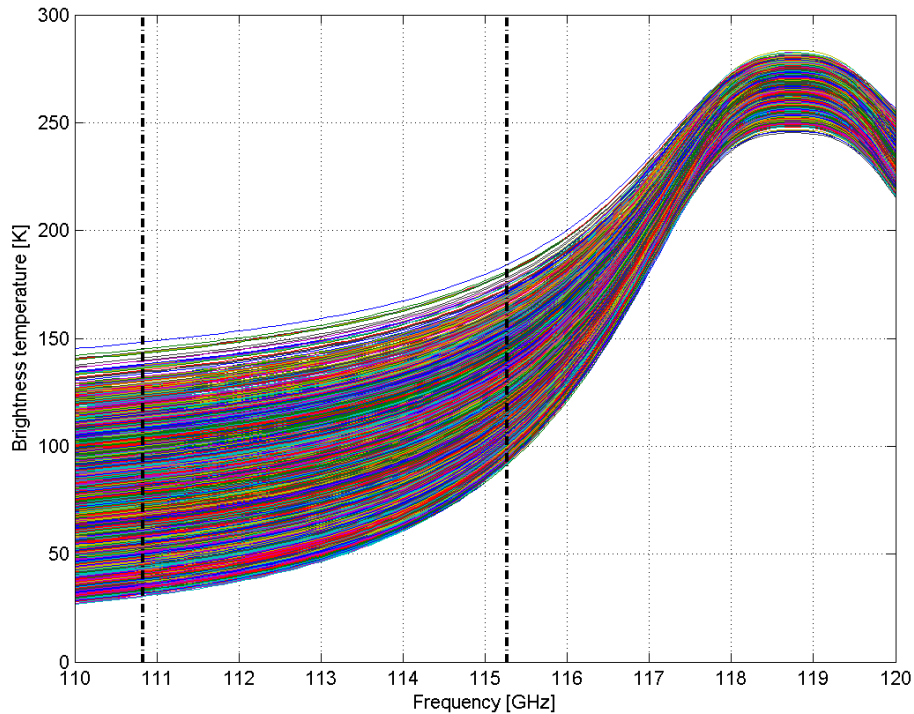


Figure 4. Simulated atmospheric spectra from one year of radiosonde data taken at the Landvetter airport 38 km NE of the Onsala site. Different tropospheric conditions explain the seen variation.

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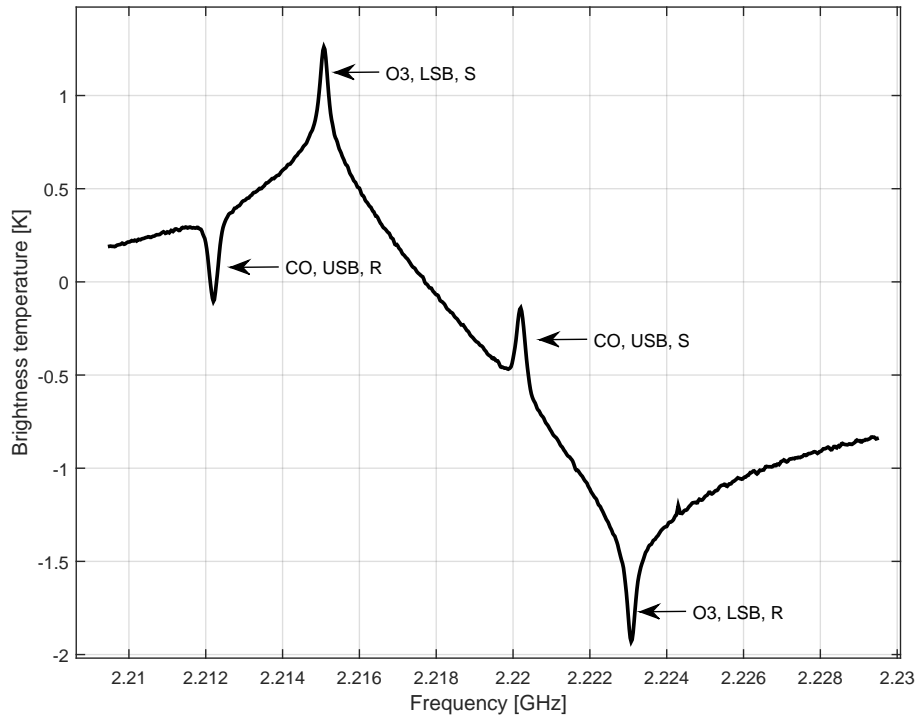


Figure 5. Average frequency-switched spectrum from December 2014 of O_3 from the lower sideband, LSB, and CO from the upper sideband, USB. S and R are the frequency-switching signal and reference phases. The wide wings of the O_3 -line explain the general baseline shape.

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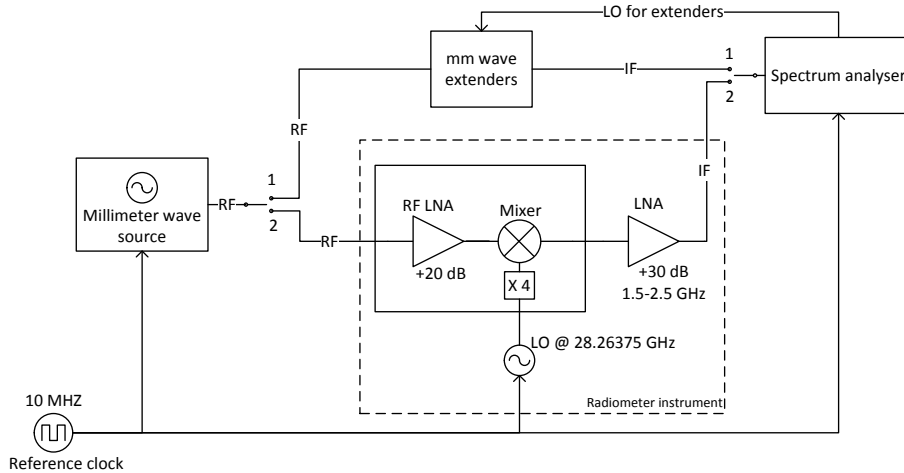


Figure 6. Setup for sideband ratio measurement.

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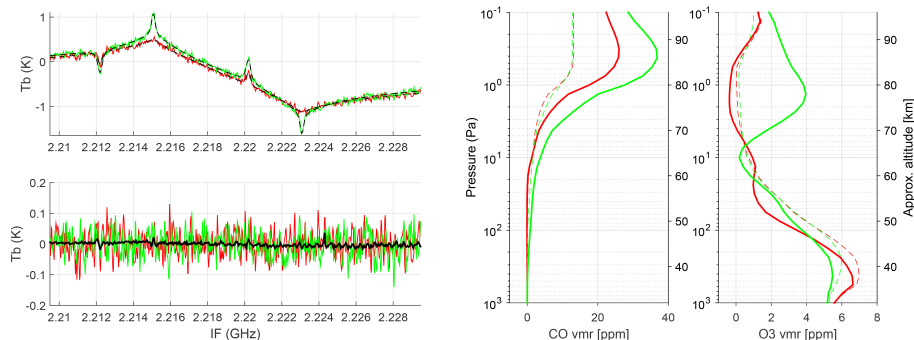


Figure 7. The top left panel shows measured spectra (solid lines) and fitted spectra (black-dashed lines) at two different times. The green line corresponds to a mid-winter night, while the red is an autumn day. The lower left panel shows the residuals from the fitting of the two spectra together with the mean residual of all spectra (black line). The two right panels show the retrieved profiles for the corresponding cases, together with the a priori profile used (dashed).

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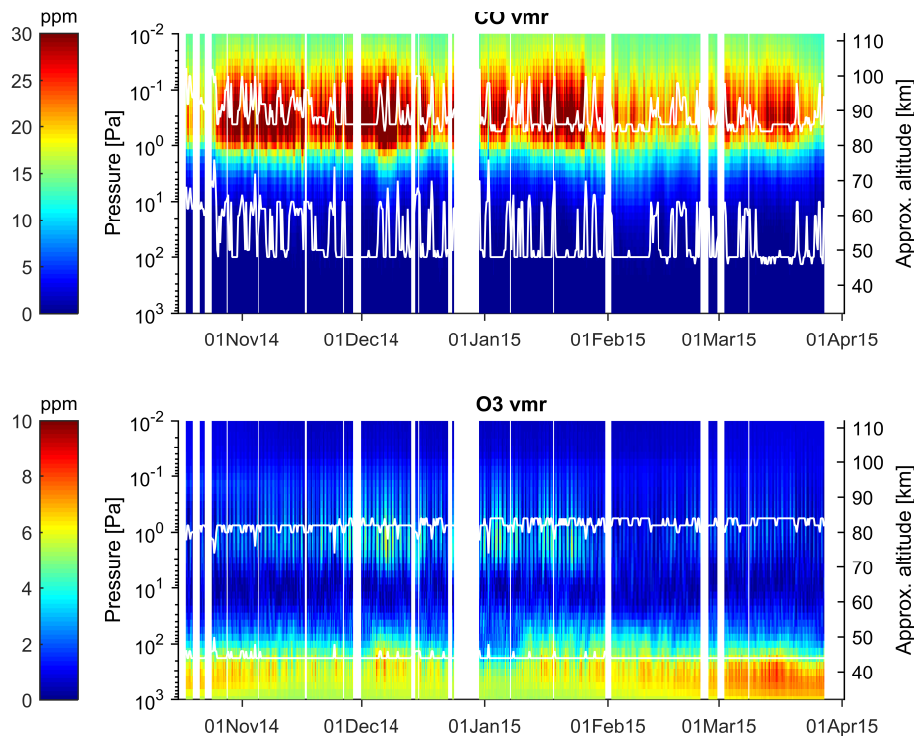
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Figure 8. Retrieved vmr of CO and O₃ (ppm) for the measurement period. The white lines mark where the a priori affects the result with 20% (< 20% between the lines).

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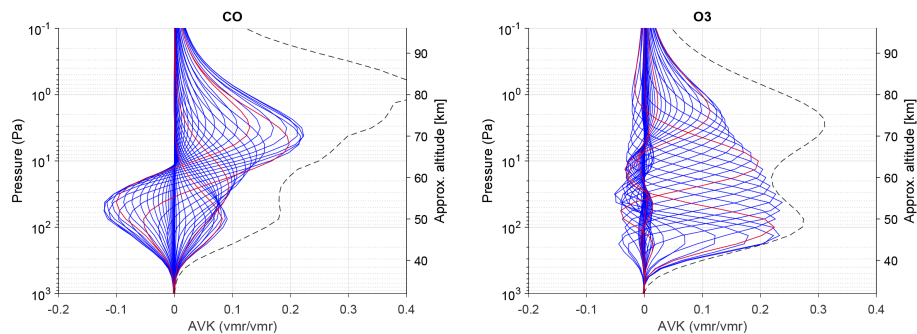


Figure 9. Averaging kernels (vmr vmr^{-1}) for CO and O_3 for 15 November 2014. The kernels for 100, 10, 1 and 0.1 Pa are highlighted with red lines. The dashed line is the measurement response divided by 4.

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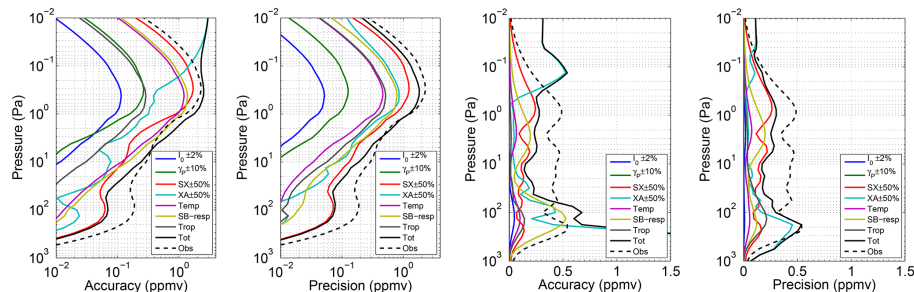


Figure 10. Accuracy and precision for CO (two leftmost panels) and O₃ (two rightmost panels) estimated by perturbation of forward model and retrieval parameters. The parameters perturbed are line strength, l_0 , pressure broadening parameter, γ_p , a priori variance, SX, a priori profile, XA, a priori temperature profile, Temp, the sideband response SB-resp and the tropospheric correction Trop. The total RMS error expected is given by the solid black line, and the observation error by the dashed black line.

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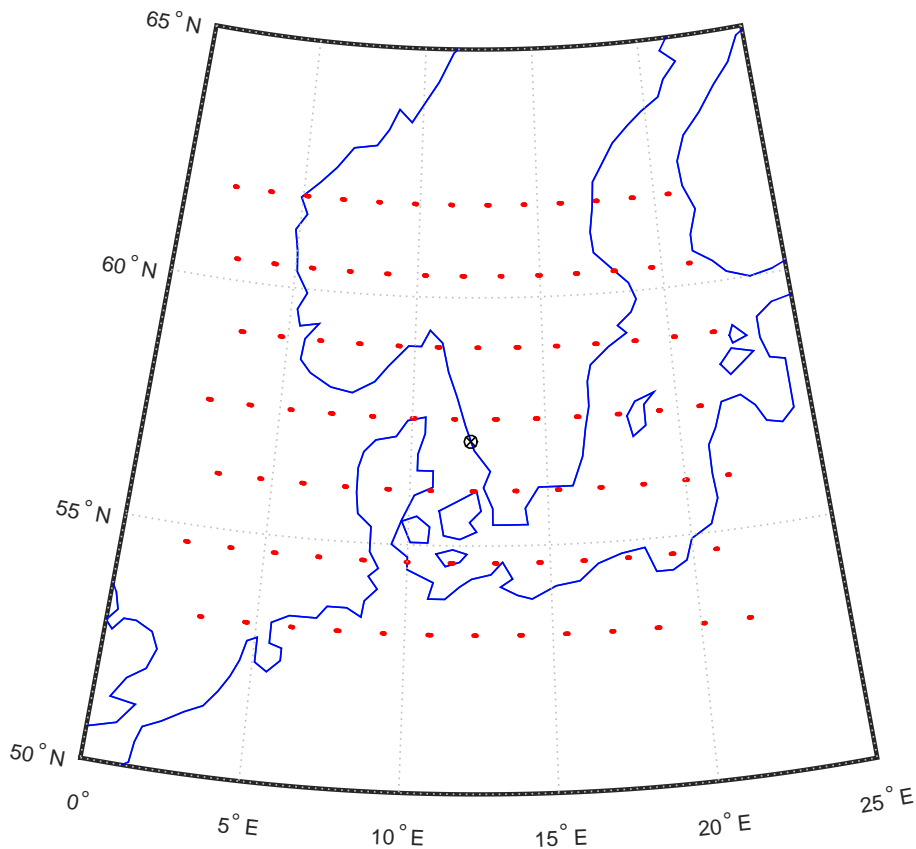


Figure 11. Collocations MLS–OSO.

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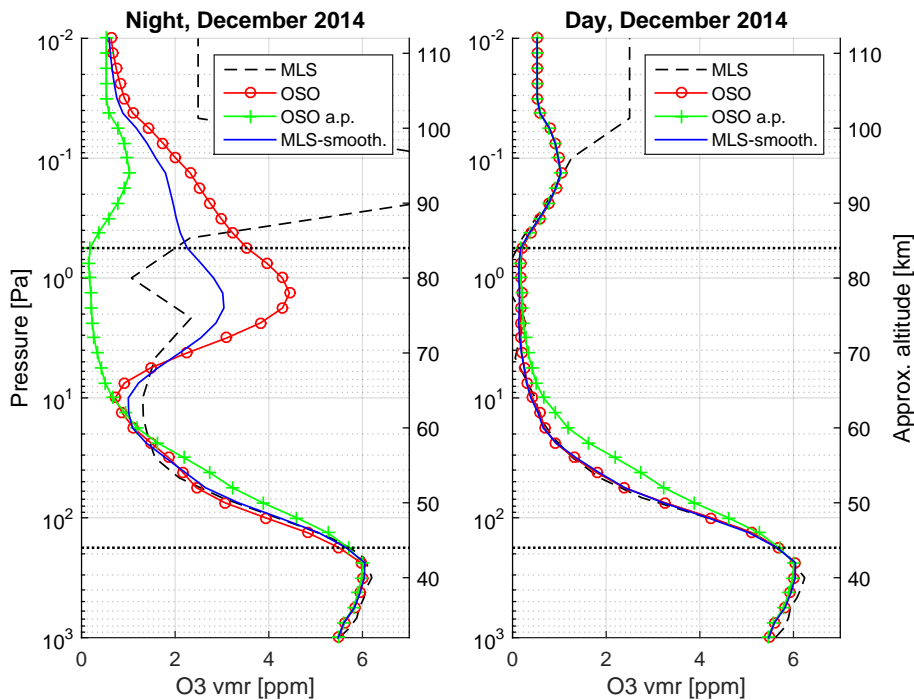


Figure 12. Average night and day vertical profiles of O₃ from December 2014. The a priori affects the result < 20% between the dotted horizontal lines.

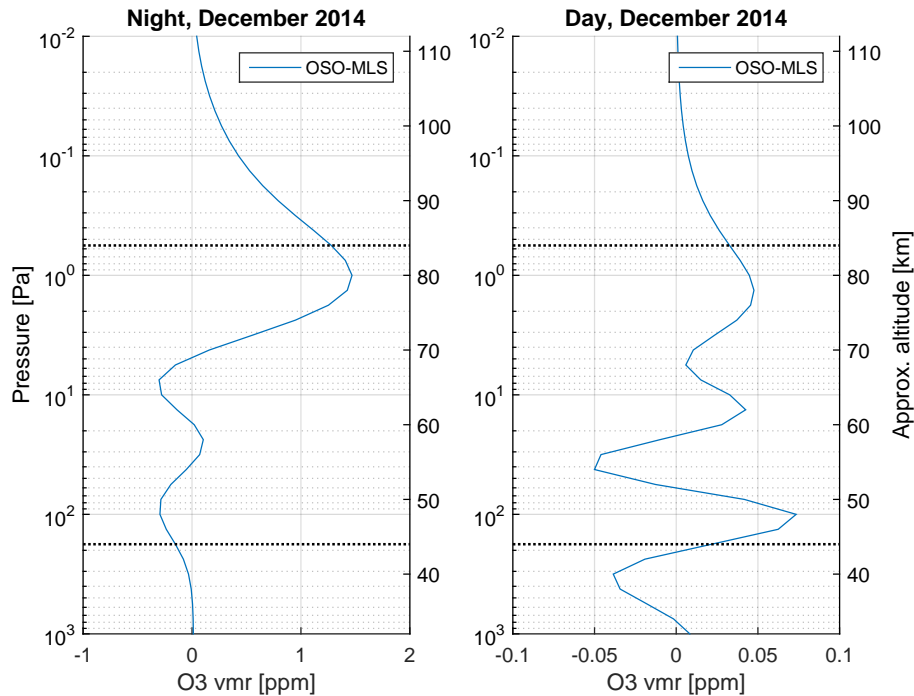


Figure 13. Difference between OSO and MLS (OSO-MLS) for night and day vertical profiles of O_3 from December 2014. The a priori affects the result $< 20\%$ between the dotted horizontal lines.

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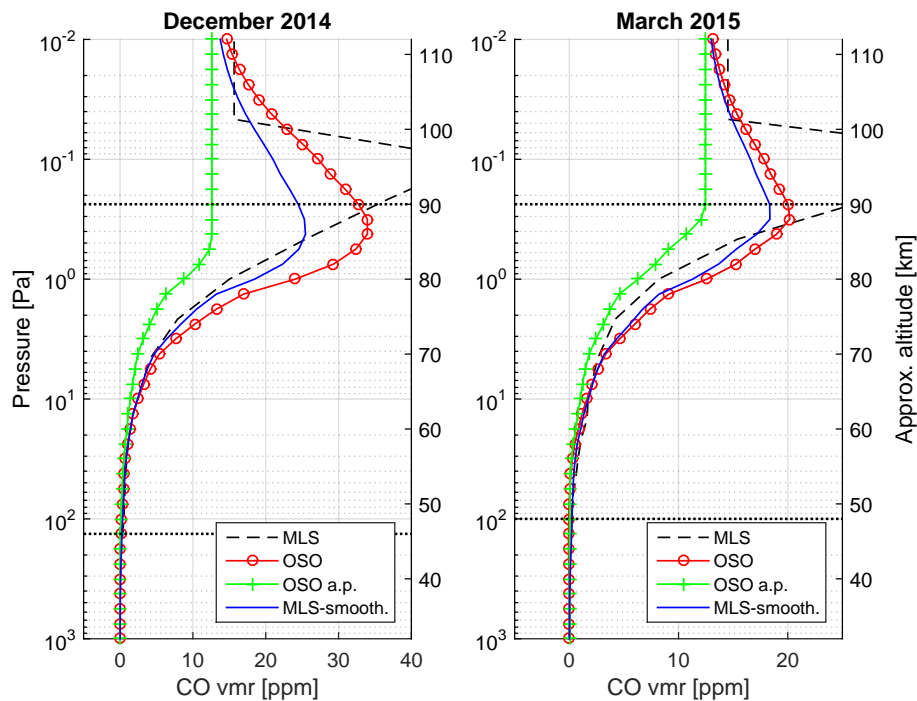


Figure 14. Average December and March vertical profiles of CO. The a priori affects the OSO result < 20% between the dotted horizontal lines.

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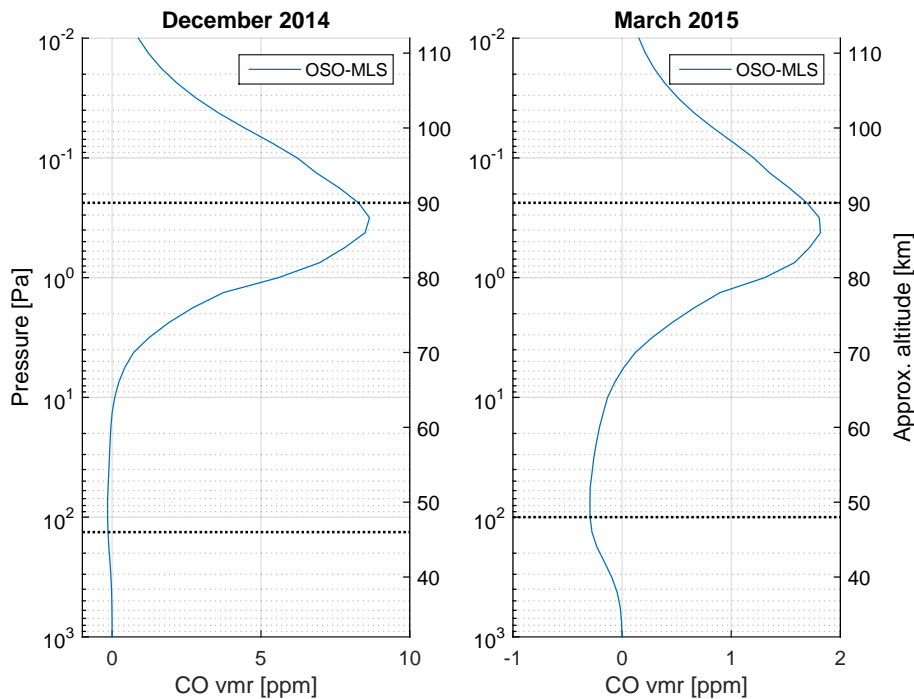


Figure 15. Difference between OSO and MLS (OSO-MLS) for December and March vertical profiles of CO. The a priori affects the result < 20% between the dotted horizontal lines.

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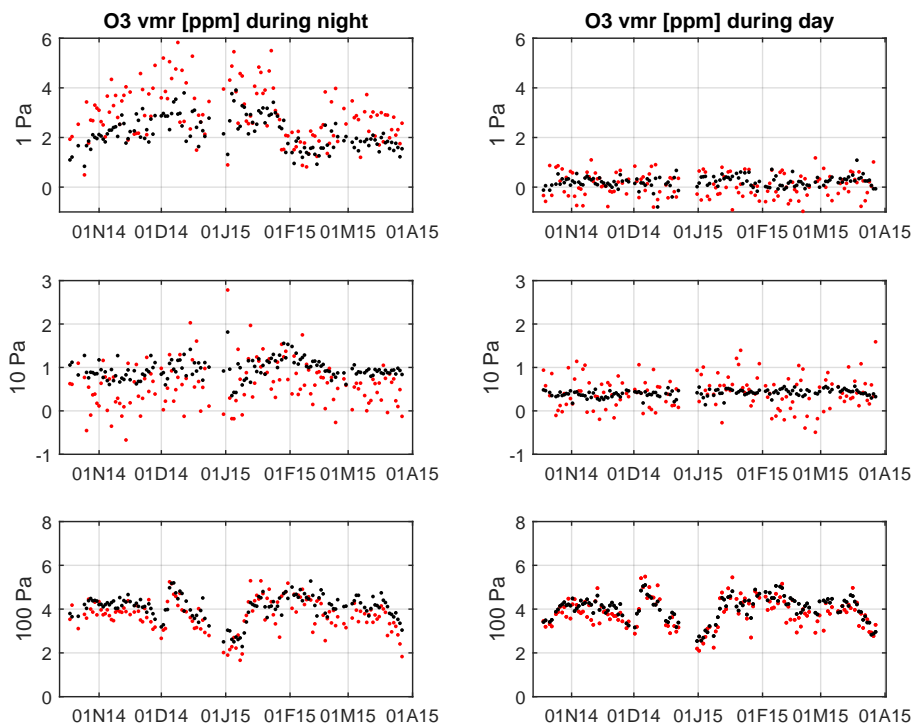


Figure 16. O₃ at 3 different altitudes, October–April (ddmyy), OSO (red), MLS (black).

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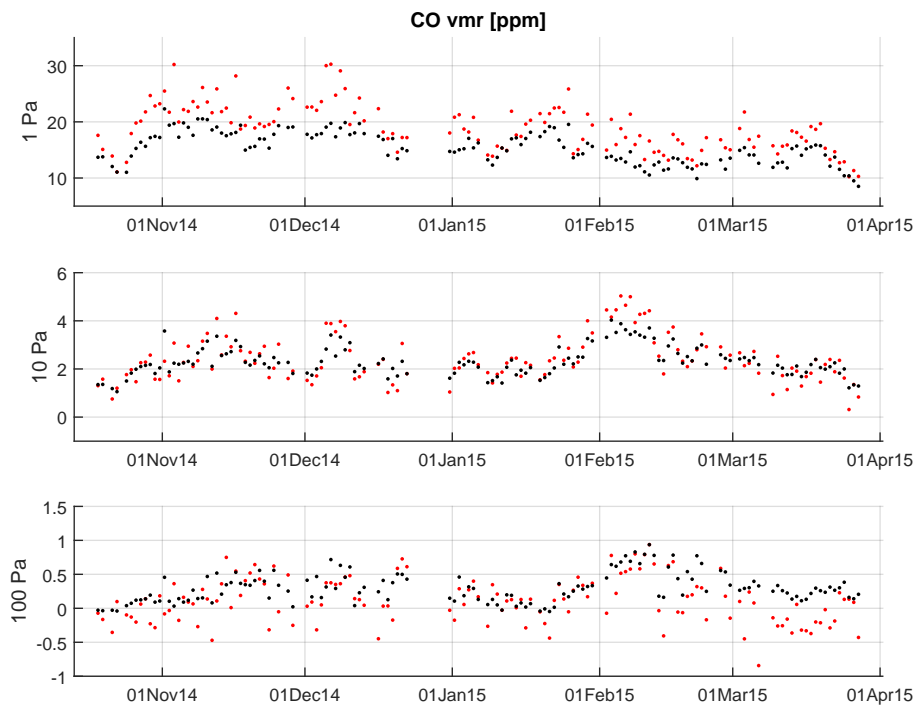


Figure 17. CO at 3 different altitudes, October–April (ddmmyy), OSO (red), MLS (black).

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