



**Arctic Snow
Microstructure
Experiment for the
development of snow
emission modelling**

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Arctic Snow Microstructure Experiment for the development of snow emission modelling

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Abstract

The Arctic Snow Microstructure Experiment (ASME_x) took place in Sodankylä, Finland in the winters of 2013–2014 and 2014–2015. Radiometric, macro-, and microstructure measurements were made under different experimental conditions of homogenous snow slabs, extracted from the natural seasonal taiga snowpack. Traditional and modern measurement techniques were used for snow macro- and microstructure observations. Radiometric measurements of the microwave emission of snow on reflector and absorber bases were made at frequencies 18.7, 21.0, 36.5, 89.0 and 150.0 GHz, for both horizontal and vertical polarizations. Two measurement configurations were used for radiometric measurements: a reflecting surface and an absorbing base beneath the snow slabs. Simulations of brightness temperatures using two microwave emission models were compared to observed brightness temperatures. RMSE and bias were calculated; with the RMSE and bias values being smallest upon an absorbing base at vertical polarization. Simulations overestimated the brightness temperatures on absorbing base cases at horizontal polarization. With the other experimental conditions, the biases were small; with the exception of the HUT model 36.5 GHz simulation, which produced an underestimation for the reflector base cases. This experiment provides a solid framework for future research on the extinction of microwave radiation inside snow.

1 Introduction

Snow is a vital component of the water cycle, as well as being critically important for meteorological and climatological studies due to its high albedo, its high thermal emissivity, and its thermal insulating properties (Cohen and Rind, 1991). In addition, over one billion people rely on snowmelt for their fresh water drinking supply (Barnett et al., 2005). To predict and monitor the evolution of potential snowmelt, continuous observations of key parameters such as Snow Water Equivalent (SWE), Snow Depth (SD),

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and Snow Extent (SE) are required throughout the year. While traditional snow pit and automatic weather station observations are important, remote sensing observations of snow with passive microwave radiometers are currently the only means in northern countries to provide vital global daily measurements of snow properties.

5 Over the last 30 years, space borne passive microwave observations have been used to estimate snow mass and SWE (Chang et al., 1987; Hollinger et al., 1990; Kelly et al., 2003; Takala et al., 2011). Numerous empirical formulae were developed for estimating SWE from the passive microwave observations (Künzi et al., 1982; Chang et al., 1987; Hallikainen, 1989; Amlien, 2008); however, relying on rigid regression coefficients, these empirical approaches were often only valid for certain regional areas with poor inter-annual consistency (Derksen et al., 2003). Purely theoretical models for snow emission have also been developed (e.g. Tsang et al., 1985). However, these models tend to be very complex and due to the diversity of ancillary information required, their use in practical SWE retrieval from satellite observations is limited. Sitting
10 between empirical and theoretical models are semi-empirical models; these combine radiative transfer theory with results from observations, adjusting key model components empirically. Two commonly used semi-empirical models are the Helsinki University of Technology (HUT) snow emission model (Pulliainen et al., 1999; Lemmetyinen et al., 2010) and the Microwave Emission Model of Layered Snowpacks (MEMLS, Wiesmann and Mätzler, 1999; Mätzler and Wiesmann, 1999).

Both the HUT snow emission model and MEMLS use snow parameters to describe the snowpack and snow microstructure. These parameters include physical temperature, density, and some form of microstructure parameter. This microstructure parameter (describing size, shape, orientation of snow grains) has a large effect on the observed brightness temperature (Foster et al., 1999; Armstrong et al., 1993) because
15 the intensity of scattered microwave radiation is directly linked to snow microstructure (Chang et al., 1987). However, the amount of scattering, described by the scattering coefficient in both the HUT snow emission model and MEMLS, is empirically defined based on observations (Pulliainen et al., 1999; Wiesmann and Mätzler, 1999). How-

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ever, MEMLS also includes an option to define the scattering coefficient purely on a physical basis (Mätzler and Wiesmann, 1999).

The Arctic Snow Microstructure Experiment (ASME_x) took place at the Arctic Research Centre of Finnish Meteorological Institute (FMI-ARC) in Sodankylä, Finland in the winter seasons of 2013–2014 and 2014–2015. During the ASME_x, macro-, microstructure and radiometric measurements of homogeneous snow slabs were made. The snow slabs were extracted from the natural seasonal taiga snowpack. The radiometric measurements were made on two different bases; one assumed perfect absorber and one perfect reflector. Observations of snow macro- and microstructure were made after radiometric measurements. The observed parameters were fed into the HUT snow emission model and MEMLS to produce simulated brightness temperatures. Only homogeneous slabs of dry snow were considered for microwave emission simulation. This was to avoid using wet snow in the radiometric measurements, as the dielectric properties of dry and wet snow are very different. The real and imaginary parts of the dielectric constant of water are much greater than those of ice (Stiles and Ulaby, 1981); increasing the complexity of the behaviour of the dielectric properties of snow.

This paper uses both HUT snow emission model and MEMLS to simulate the microwave emission of homogenous snow slabs extracted from the natural snowpack in FMI-ARC during ASME_x, and compares simulated and observed microwave emission from the snow slabs. The ultimate aim of ASME_x is to improve the understanding of the microwave extinction processes within the snowpack, and their relation to microstructural properties of natural snow cover. This will enable to improve the precision of future and existing snow emission models.

2 Methods and models

2.1 Methods

2.1.1 Excavation of snow slabs

The snow slabs were extracted from the natural snowpack in the Intense Observation Area (IOA) of the FMI-ARC that is situated in the clearing of a sparse pine forest, in So-

dankylä, Finland. A snow sample of size 80 cm × 60 cm was mechanically removed from the snowpack. The thickness of these slabs varied typically between 14 and 19 cm, with the exception of one slab, comprised of depth hoar, being approximately 5 cm thick. The preparation and extraction of the snow slabs was delicate process.

Snow slabs were taken from different depths in order to capture a range of grain sizes and types. The snow slabs were taken at different periods during the two winters of ASME_x; to capture a wider range of grain sizes and types. Each of the snow slabs was extracted from a homogeneous layer, and its stratigraphy was manually assessed after the radiometric measurements. Once a homogeneous layer of sufficient thickness was selected, the sample was prepared by pushing a metal plate (surrounded by a microwave transparent plastic sheet (transparent to microwaves) to avoid snow freezing to the metal plate) into the snowpack and selecting the snow sample with a plastic frame as shown in Fig. 1. The top of the snow sample was carefully smoothed with a metal plate as gently as possible without making artificial features on the slab surface. Immediately after extraction, the slabs were placed in front of the radiometer for brightness temperature measurements. 14 samples were extracted in that manner.

2.1.2 Radiometric measurements

The microwave radiometric measurements were made with two RPG-XCH-DP Dicke Switch radiometers, installed on top of the radiometric tower in the IOA. The experimental set up of radiometric measurements is described in Fig. 2. Five different fre-

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quencies (18.7, 21.0, 36.5, 89.0, and 150.0 GHz) at both horizontal and vertical polarizations were used; although not all frequencies were working for all slabs. Tables 1 and 2 detail the radiometric data collected from the ASME_x slabs in 2014 and 2015 respectively. The radiometric measurements were made at an inclination angle of 50°.

Equivalent sky brightness temperatures were also made to measure the intensity of the downwelling radiation.

Radiometric measurements followed a comparable procedure as in Wiesmann et al. (1998). The first measurement was made with the snow slab on top of the reflective metal base. The metal base acts as a perfect reflector by reflecting the downwelling emission of microwave radiation from the sky. Once the snow slab had been observed at all frequencies, sky measurements at an equivalent incidence angle were made. The metal plate was then carefully removed from the set up, so that the snow slab was upon the assumed perfect absorber. The radiometric measurements were then repeated. Emissivity tests of the absorbing material, using the experimental setup in Fig. 2 without the snow slab and metal plate, proved that the assumption of a near-perfect blackbody was valid for all slab experiments, with the exception of slabs B05 and B07. For these two slabs, the metal strips in the tape used to hold the top-most piece of Styrofoam together caused a reduction in brightness temperature at horizontal polarizations, at different frequencies. A correction (none at 18.7 GHz, -1 K at 21.0 GHz, -2 K at 36.5 GHz, -8 K at 89.0 GHz, and -15 K at 150.0 GHz) was applied to the absorbing base brightness temperature data for slabs B05 and B07. Throughout the radiometric measurements the physical temperatures of snow, air, and absorbing material were also measured for modelling purposes.

2.1.3 Measurements of snow macro- and microstructure

Once the radiometric measurements had been completed, the destructive sampling of the physical parameters of snow macro- and microstructure took place. Initially, the stratigraphy of the slab was observed using the SnowMicroPen (SMP; Schneebeli and Johnson, 1998; Schneebeli et al., 1999). The SMP uses a sensitive piezoelectric force

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sensor on top of a penetrative rod, which is capable of detecting changes in penetrative resistance at a high resolution ($4\ \mu\text{m}$). 12 SMP profiles were taken across each slab to assess the stratigraphy and homogeneity. The nominal locations of the SMP profiles, as well as all other macro- and microstructure measurements, are shown in Fig. 3.

From the SMP profiles, it is possible to detect layers as well as produce profiles of density, correlation length, and Specific Surface Area (SSA) of the snow (Proksch et al., 2015). For the purposes of this experiment, nine slabs (eight dry and one wet) could be considered homogeneous, with minimal horizontal and vertical features. The other five slabs exhibited significant vertically layered structures, and contained features such as ice crusts within the snow. These internal features would produce additional scattering and internal reflections; that would be difficult to quantify in the models for simulation.

The SSA is defined as the ratio between the surface area of the ice and its mass (Legagneux et al., 2002). A new method for measuring the SSA is the IceCube instrument, which is a commercially available version of DUFISSS (Gallet et al., 2009). The IceCube instrument uses a 1310 nm infrared laser to measure the hemispherical reflectance from the sample. The SSA of the snow slab was measured at two different locations in a vertical profile with 3 cm intervals. Traditional grain size, E (Fierz et al., 2009), macro-photographs were taken from every SSA sample (Leppänen et al., 2015) in order to have profiles of traditional grain size in the same locations as the SSA profiles.

Snow samples were taken from the centre of the radiometer footprint to be scanned with micro-computed tomography (micro-CT) apparatus. The cast samples were analysed via three-dimensional x-ray tomography in WSL Institute of Snow and Avalanche Research, SLF, Switzerland, to produce a three dimensional image of the snow (Heggli et al., 2009). From this image, it is possible to measure many important microstructural parameters, especially a vertically highly resolved profile of density and correlation length.

In addition to the different number of microstructural measurements of the snow slab, vertical profiles of physical temperature and density took place in other locations

within the slab with a vertical resolution of 5 cm. The density profiles were made using a density cutter with a volume of 500 cm³.

2.2 Models

2.2.1 Helsinki University of Technology snow emission model

5 The Helsinki University of Technology (HUT) snow emission model (Pulliainen et al., 1999; Lemmetyinen et al., 2010) is a semi-empirical model, which uses the radiative transfer approach to model the microwave brightness temperature. It is capable of treating the snow as a single homogeneous layer (Pulliainen et al., 1999) or as a series of homogeneous layers (Lemmetyinen et al., 2010), with the layers being defined by its
10 physical temperature, density, observed grain diameter, and SWE.

The model's basic assumption is that the microwave radiation is scattered mostly in the forward direction, which allows simplifying the radiative transfer equation to a single flux. Calculation of the absorption coefficient in the HUT model is based on empirical models by Mätzler (1987); while the total extinction coefficient (sum of absorption and
15 scattering coefficients) was originally calculated by Hallikainen et al. (1987) from observations of snow in northern Finland. Other extinction coefficient relationships exist for the HUT model (Roy et al., 2004; Kontu and Pulliainen, 2010), but these have not been used here.

20 The HUT model uses up- and downwelling emissions, represented by single-flux approximations, to calculate the total emission at the top of the snowpack. Multiple reflections at layer interfaces are accounted. Separate modules were used to simulate the effect of vegetation and atmosphere to detected emission were published with the original model, but were not applied here.

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2.2.2 Microwave Emission Model of Layered Snowpacks

MEMLS (Wiesmann and Mätzler 1999; Mätzler and Wiesmann, 1999) is also based on radiative transfer theory, treating the snowpack as a stack of horizontal layers, with each layer being characterised by its depth, physical temperature, density, and exponential correlation length. These data are used to calculate absorption and scattering coefficients within the snow, as well as transmissivity and reflectivity between adjacent snow layers. A two flux (up and down welling) model is used to calculate the emitted brightness temperature at the top of the snowpack. However, the absorption and scattering coefficients are adjusted with six flux coefficients (up- and downwelling, and four horizontal directions). The scattering coefficient was empirically defined from radiometric and macro- and microstructure measurements as laid out by Wiesmann et al. (1998). An optional feature, originally implemented for coarse-grained snow with a large correlation length, is to use the Improved Born approximation (Mätzler and Wiesmann, 1999) for the calculation of the scattering coefficient.

3 Results

3.1 Macro- and microstructure observations

Preliminary analyses of snow macro- and microstructure measurements include slab depth, physical temperature, density, SSA, grain size E and homogeneity. E ranged from 0.5 to 2.0 mm, average slab densities ranged from 130 to 340 kg m⁻³, and physical snow temperature ranges from -14 to 0 °C. The variability of snow characteristics and homogeneity from all 14 slabs are shown in Table 3. Final results from SMP and micro-CT measurements will be used in future research for the quantification of snow microstructure.

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3.2 Comparison of snow emission models

The parameters from the eight dry homogeneous slabs in Table 3 were fed into both the single layer HUT snow emission model and into MEMLS to produce simulated brightness temperatures. The ground layer in both of the models was modified to simulate the absorbing and reflecting bases by altering the reflecting properties of the ground, to model the reflective properties of the metal plate ($r = 1$) and the absorbing base ($r = 0$). The absorbing and reflective bases were simulated assuming a near-perfect absorption and reflection at the snow–base interface. The directly measured downwelling sky contribution was applied as the downwelling flux in both models. The simulated brightness temperatures were compared to the observed brightness temperatures, and the RMSE and bias were calculated for 18.7, 21.0, and 36.5 GHz at both horizontal and vertical polarizations. The RMSE and bias at the two higher frequencies (89.0 and 150.0 GHz) were not calculated for this study. Figure 4 shows the RMSE values and Fig. 5 shows the bias values of the simulations.

The values in Fig. 4 show that for the absorbing base, the HUT model simulations tend to have smaller RMSE values than MEMLS; while for the reflective base simulations the RMSE values are comparable at 18.7 and 21.0 GHz. At 36.5 GHz the HUT snow emission model produces larger RMSE values than MEMLS. The RMSE values for the absorbing base of vertical polarization (V-ABS) are the smallest.

It can be seen from Fig. 5 that the reflective base cases have the smallest bias, with 18.7 and 21.0 GHz only having very small magnitude ($< |2\text{ K}|$) biases. At 36.5 GHz, the HUT model is negatively biased and MEMLS is slightly positively biased. It suggests that the HUT model underestimated the microwave emission while MEMLS slightly overestimated it at 36.5 GHz on the reflective base cases. The bias for the absorbing base of horizontal polarization (H-ABS) was positive regardless of model or frequency. The bias for the V-ABS simulations ($< |6\text{ K}|$) was negative for all frequencies with MEMLS, and with HUT model positive at 18.7 and 21.0 GHz, and slightly negative at 36.5 GHz.

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One potential reason for the H-ABS large positive bias and larger RMSE values is an underestimation in scattering, potentially caused by slight changes in the snow density and microstructure parameter that were not recorded by the traditional observations, due to the coarse resolution of the method (vertical profile of 3–5 cm). By measuring the snow density or microstructure parameter through the use of modern techniques (such as the SnowMicroPen or micro-CT), the bulk average of these parameters would be more accurate, and could lead an improvement in H-ABS.

4 Summary

The Arctic Snow Microstructure Experiment (ASME_x) consisted of radiometric, macro- and microstructure measurements of snow slabs upon absorbing and reflecting bases. Brightness temperatures of the homogeneous snow slabs were simulated with the HUT snow emission model and with MEMLS. Results of the comparison of simulations and observations are described in Sect. 3.2. The HUT model produced smaller RMSE across all three frequencies for the simulations upon an absorbing base. The reflective base simulations produced RMSE values that were comparable with the HUT model and MEMLS at 18.7 and 21.0 GHz. Both models overestimated the brightness temperature at H-ABS, and at V-ABS the single layer HUT model slightly overestimated the brightness temperature while MEMLS underestimated it. Both models produced very small biases for the reflective base cases, with the exception of the HUT model at 36.5 GHz.

The RMSE and bias is produced by internal extinction processes within the snow slabs, which are not captured by the model input data, or which are imperfectly simulated by the model physics. The relatively high errors especially at H-pol, considering the highly controlled measurement setup, highlight the requirement for further development of the models, as well as the need to better quantify the snow microstructural properties themselves. These preliminary brightness temperature simulations will be repeated in the future using the physical snow properties collected by the modern

techniques including SMP and micro-CT measurements. Ultimately, a revised extinction model will be created for the HUT snow emission model, and implemented with the aim to improve the model inversions of SWE from radiometric measurements of microwave emission.

5 *Author contributions.* J. Lemmetyinen, M. Sandells, and M. Schneebeli planned the experiment with help of W. Maslanka. W. Maslanka had main responsibility in organizing and carrying out the measurements. L. Leppänen had responsibility of organizing measurements as local operator including implementation of measurement setup and participation to measurement procedure. A. Kontu and H.-R. Hannula participated to experimental measurements. W. Maslanka
10 made data processing and simulations. W. Maslanka and L. Leppänen prepared the manuscript with contributions from all co-authors.

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Table 1. Radiometric data measured from the 2014 ASME_x slabs (Horizontal/Vertical Polarization). Brightness temperatures from reflective base ($T_{B,M}$), sky after reflective base measurements ($T_{B,SKY,M}$), absorbing base ($T_{B,A}$) and sky after absorbing base measurements ($T_{B,SKY,A}$) are presented.

Sample	Frequency (GHz)	$T_{B,M}$ (K)	$T_{B,SKY,M}$ (K)	$T_{B,A}$ (K)	$T_{B,SKY,A}$ (K)
A01	18.7	22.94/23.43	18.06/17.90	243.27/258.29	14.01/13.24
A01	21.0	30.42/30.87	24.11/24.01	238.51/257.81	18.59/18.47
A01	36.5	51.66/52.13	38.21/38.58	243.81/257.82	25.23/25.51
A02	21.0	46.41/54.53	15.22/15.83	152.52/217.99	14.39/14.85
A02	36.5	93.71/115.81	23.89/24.98	144.62/192.49	22.21/23.29
A03	18.7	37.28/35.96	19.57/19.83	255.18/272.08	19.57/19.83
A03	21.0	50.94/50.88	32.01/32.13	257.10/272.19	32.01/32.13
A04	18.7	77.45/79.14	19.30/19.16	221.22/254.11	19.30/19.16
A04	21.0	111.81/111.96	29.17/29.09	217.65/248.46	29.17/29.09
A05	18.7	47.95/48.96	19.43/17.84	225.29/262.26	19.79/18.51
A05	21.0	63.11/64.68	28.63/28.70	239.14/263.03	29.36/29.46
A05	89.0	180.79/201.66	54.13/53.37	193.85/202.99	65.79/65.02
A05	150.0	205.16/212.83	96.23/100.98	205.65/212.11	109.70/111.87
A06	18.7	28.54/31.80	7.40/9.11	228.19/256.50	8.29/9.89
A06	21.0	41.92/45.64	12.63/12.23	235.55/256.59	13.46/12.77
A06	89.0	177.19/192.06	27.48/27.36	185.05/198.34	29.58/27.76
A06	150.0	181.88/188.24	42.61/34.88	184.20/189.46	47.67/35.22
A07	18.7	27.34/27.94	9.78/10.20	229.06/258.48	10.59/11.00
A07	21.0	37.85/38.03	13.80/14.40	224.12/257.64	14.55/14.59
A07	89.0	165.43/183.30	30.17/30.54	168.96/186.30	30.17/30.54
A07	150.0	175.07/186.75	45.36/39.93	174.46/187.88	45.36/39.93

Table 2. Radiometric data measured from the 2015 ASMEEx slabs (Horizontal/Vertical Polarization). Brightness temperatures from reflective base ($T_{B,M}$), sky after reflective base measurements ($T_{B,SKY,M}$), absorbing base ($T_{B,A}$) and sky after absorbing base measurements ($T_{B,SKY,A}$) are presented.

Sample	Frequency (GHz)	$T_{B,M}$ (K)	$T_{B,SKY,M}$ (K)	$T_{B,A}$ (K)	$T_{B,SKY,A}$ (K)
B01	18.7	18.73/18.33	10.87/9.81	246.71/257.23	11.04/9.74
B01	21.0	28.69/28.30	17.92/17.49	245.98/258.02	17.83/17.46
B01	36.5	42.71/44.08	21.99/22.48	241.77/254.49	21.58/21.81
B02	18.7	20.60/19.30	8.66/8.21	229.52/251.64	7.51/6.49
B02	21.0	30.56/28.61	13.69/13.72	231.56/252.80	11.77/11.65
B02	36.5	54.75/54.55	24.00/24.27	228.46/247.58	19.49/19.86
B03	18.7	24.64/23.74	13.33/11.84	234.63/264.67	13.35/11.48
B03	21.0	37.70/36.85	23.79/23.76	243.67/265.30	23.15/23.13
B03	36.5	57.32/56.63	25.16/26.32	241.55/263.11	24.26/25.41
B04	18.7	24.78/22.95	10.38/9.12	229.91/261.78	9.64/9.03
B04	21.0	31.14/30.06	13.61/13.61	232.52/260.89	13.30/11.48
B04	36.5	64.69/63.53	21.54/21.80	229.18/255.05	21.21/21.67
B05	18.7	35.56/33.28	7.71/8.02	233.71/255.74	8.01/8.29
B05	21.0	43.27/41.56	13.73/12.75	242.98/257.92	14.28/13.40
B05	36.5	85.96/88.68	20.57/21.87	226.34/241.98	20.84/21.94
B05	89.0	164.30/163.60	36.50/35.80	180.20/177.10	34.40/37.90
B05	150.0	207.00/191.90	53.30/54.90	203.00/190.40	53.60/58.30
B06	18.7	22.89/22.31	9.02/8.43	238.45/260.81	9.11/8.27
B06	21.0	30.09/28.78	12.46/12.46	239.06/261.53	11.49/11.37
B06	36.5	63.62/62.54	19.83/21.34	236.89/258.01	19.95/21.34
B06	89.0	195.41/200.85	28.96/30.05	208.27/210.84	30.43/31.09
B06	150.0	201.43/194.45	47.82/44.46	202.49/193.59	45.18/43.68
B07	18.7	55.39/55.25	10.18/9.65	209.77/242.70	10.08/9.50
B07	21.0	62.84/64.10	11.72/12.22	214.42/242.58	11.23/11.74
B07	36.5	124.04/131.10	19.12/20.80	195.59/211.41	19.83/20.91
B07	89.0	167.92/165.34	29.24/29.18	167.85/165.21	27.02/27.93
B07	150.0	190.38/180.01	42.67/45.89	189.56/179.56	36.66/44.31

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Table 3. Averaged results from macro- and microstructure measurements. It should be noted that slab A03 was wet, so was not considered for model simulation.

Date	Slab Reference	Temperature (°C)	Density (kg m ⁻³)	Grain size (mm)	SSA (m ² kg ⁻¹)	Depth (cm)	Homogenous
13 Jan 2014	A01	-13.1	138	0.5	35.8	17.8	Yes
14 Jan 2014	A02	-22.2	269	0.7	15.4	15.6	No
11 Feb 2014	A03	-0.3	228	0.6	18.0	16.6	Yes
13 Feb 2014	A04	-0.5	226	0.9	11.3	18.0	No
3 Mar 2014	A05	-0.8	287	0.9	15.8	15.6	No
18 Mar 2014	A06	-7.6	280	0.8	17.5	14.8	Yes
20 Mar 2014	A07	-5.1	285	0.9	15.5	14.8	No
2 Feb 2015	B01	-13.2	140	0.5	36.4	14.8	Yes
5 Feb 2015	B02	-10.9	160	0.5	36.4	13.9	Yes
19 Feb 2015	B03	-2.6	234	0.6	22.8	14.9	Yes
11 Mar 2015	B04	-5.4	268	1.1	21.2	16.2	Yes
12 Mar 2015	B05	-3.2	337	1.9	10.3	5.4	Yes
24 Mar 2015	B06	-5.4	317	1.3	17.3	14.5	Yes
25 Mar 2015	B07	-3.7	283	2.0	9.4	15.2	No



Figure 1. Snow sample was taken from snowpack with a plastic frame and a metal bottom plate surrounded by a plastic sheet.

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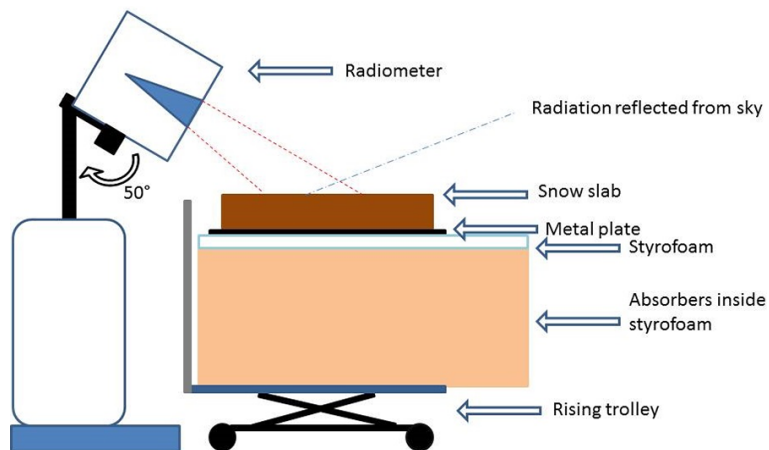


Figure 2. Setup for radiometer measurements with 50° inclination angle.

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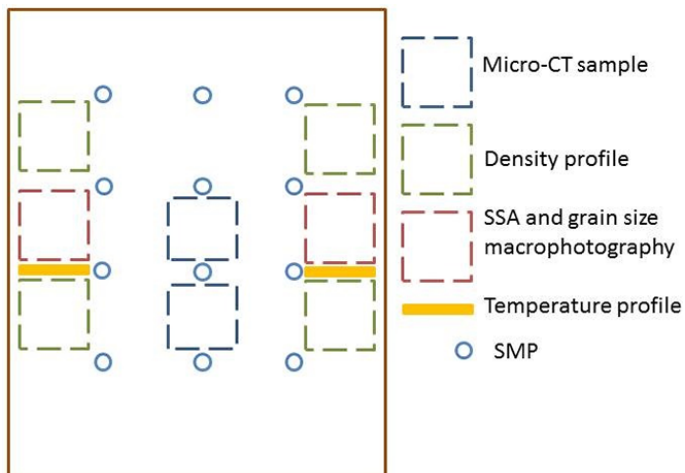


Figure 3. Approximate location of the macro- and microstructure measurements in the snow slab.

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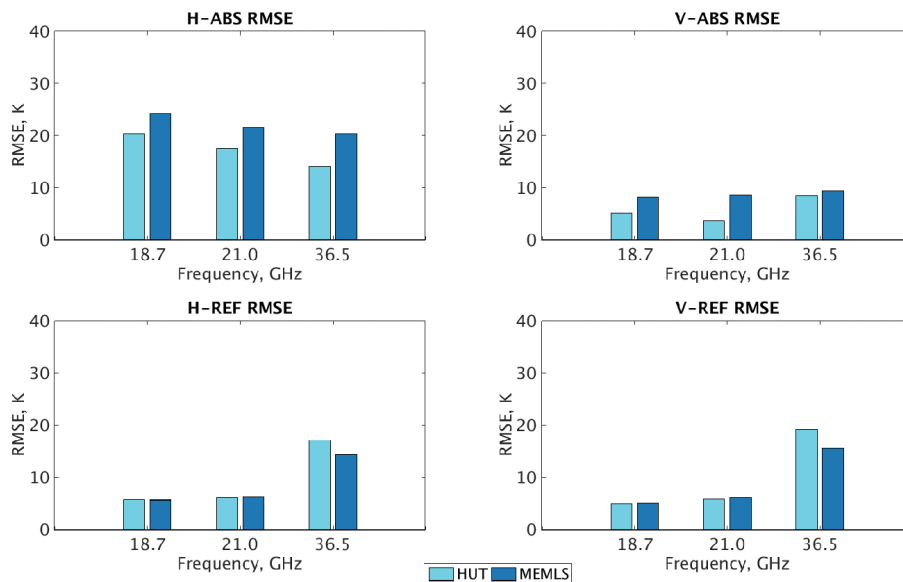


Figure 4. Simulated brightness temperature RMSE at horizontal (H) and vertical (V) polarizations for the absorber material base (ABS) and the reflective metal plate base (REF). Eight slabs were simulated at 18.7 and 21.0 GHz, while seven slabs were simulated at 36.5 GHz.

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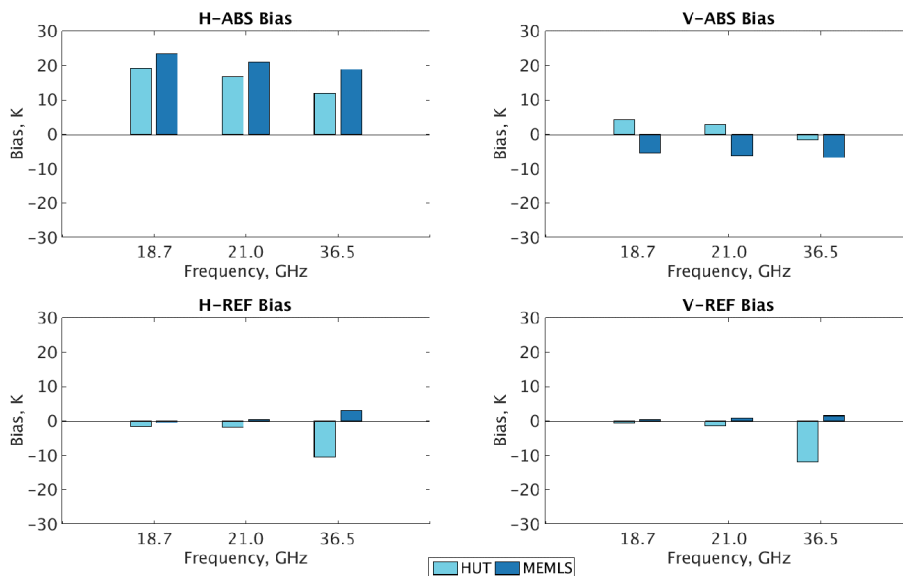


Figure 5. Simulated brightness temperature bias at horizontal (H) and vertical (V) polarizations for the absorber material base (ABS) and the reflective metal plate base (REF). Eight slabs were simulated at 18.7 and 21.0 GHz, while seven slabs were simulated at 36.5 GHz.