

1 **Data flow of spectral UV measurements at Sodankylä and** 2 **Jokioinen**

3
4 **J.S. Mäkelä^{1*}, K. Lakkala^{1,2}, T. Koskela^{1**}, T. Karppinen², J.M. Karhu², V. Savastiouk⁴,**
5 **H. Suokanerva², J. Kaurola¹, A. Arola³, A.V. Lindfors¹, O. Meinander¹, G. Leeuw^{1,5} and**
6 **A. Heikkilä¹**

7 [1] {Finnish Meteorological Institute, Research and Development, 00101 Helsinki, Finland}

8 [2] {Finnish Meteorological Institute, Sodankylä, Finland}

9 [3] {Finnish Meteorological Institute, Kuopio, Finland}

10 [4] {IOS Inc, Toronto, Canada}

11 [5] {Department of Physics, University of Helsinki, 00014 Helsinki Finland}

12 [*]{now at University of Jyväskylä, Jyväskylä, Finland}

13 [**]{now an independent researcher}

14
15 Correspondence to: K. Lakkala (Kaisa.Lakkala@fmi.fi)

16 **Abstract**

17 The data flow involved in a long-term continuous solar spectral UV irradiance monitoring
18 program is investigated and structured to provide an overall view on the multi-phase process
19 from data acquisition to the final products. The program employing Brewer
20 spectrophotometers as measuring instruments is maintained by the Finnish Meteorological
21 Institute (FMI) ever since the 1990's at two sites in Finland, Sodankylä (67°N) and Jokioinen
22 (61°N). It is built upon rigorous operation routines, processing procedures, and tools for
23 quality control (QC) and quality analysis (QA) under continuous development and evaluation.
24 Three distinct levels of data emerge, each after certain phase in the data flow: Level 0
25 denoting raw data, Level 1 meaning calibrated data processed near-real-time, and Level 2
26 comprising of post-processed data corrected for all distinguishable errors and known
27 inaccuracies. The final products disseminated to the users are demonstrated to result from a

28 process with a multitude of separate steps, each required in the production of high quality data
29 on solar UV data radiation at the Earth's surface.

30

31 1 **Introduction**

32

33 The Brewer spectrophotometer (Brewer) was initially designed to measure total column
34 ozone with the differential absorption method (Bais et al. 1996; Brewer, 1973) using the
35 direct sun port. In addition, a global (diffuser) port was introduced for measurements of
36 spectral UV. The direct sun data have also been used to calculate SO₂ (Cappellani and Bielli,
37 1995), aerosol optical depth (Gröbner et al. 2001; Kazadzis et al. 2005; Marenco et al. 2002)
38 and NO₂ (e.g. Cede et al. 2006; Diémoz et al. 2014). At present, there are over 220
39 instruments set up by research institutes all over the world (<http://kipazonen-brewer.com/>).
40 These instruments form an important network for monitoring changes in the atmospheric total
41 ozone column and solar UV irradiance at the Earth's surface.

42

43 The first Brewer spectrophotometers in Finland were set up in 1990 and 1995, in Sodankylä
44 and Jokioinen, respectively, to respond to the need to monitor total ozone and UV radiation
45 after the discovery of the Arctic ozone loss. These time series of solar spectral UV irradiance
46 of over twenty years are unique and among the longest measured in the Arctic. The
47 homogenized time series have been used for several studies related to Arctic ozone loss (e.g.,
48 Bernhard et al. 2013; Manney et al. 2011; Knudsen et al. 1998), and validation of satellite
49 data (Hassinen et al. 2008). They have also been used in, e.g., studies on the effects of UV
50 radiation on biological objects (e.g., Lappalainen et al. 2010; Martz et al. 2009), materials
51 (Heikkilä 2014) and human health (Kazantzidis et al. 2009).

52

53 The Brewer spectrophotometer is a versatile but also a complex instrument in comparison to,
54 for instance, broadband UV meters. Many intermediate steps and corrections are needed in the
55 processing of the data. The high dynamical range of five to six orders of magnitude of UV
56 irradiance reaching the surface of the Earth puts great demands to the instruments designed to
57 monitor both the short UV-B wavelengths (290-315 nm) and the longer UV-A wavelengths
58 (315-400 nm). The challenge is to maintain the sensitivity of the instrument at all
59 wavelengths. The location of Finland at high latitude, where high solar zenith angles (SZA)
60 are frequent, brings additional challenges, as the weak signal at UV-B wavelengths is near the

61 noise level of the instrument. High quality data can only be ensured after careful
62 characterization of the instrument, correction of known measurement errors and careful
63 quality control (QC) and quality assurance (QA). (Seckmeyer et al. 2001; Garane et al. 2006;
64 Lakkala et al. 2008; Webb et al. 2003).

65
66 Maintaining a Brewer spectrophotometer absolutely calibrated is a demanding task (Bernhard
67 and Seckmeyer 1999, Webb et al. 1998). International campaigns are regularly organized to
68 evaluate the calibration and measurement procedures performed by different Brewers and
69 institutes. The difficulty of the absolute calibration was already seen in the intercomparison
70 campaigns of the 1990's (Josefsson et al. 1994, Koskela et al. 1997) and in the twenty-first
71 century (e.g., Bais et al. 2001), in which the range of the deviations from the reference for UV
72 spectra was up to $\pm 20\%$. Despite the efforts to homogenize the measurements, in the last
73 European Brewer comparison organized by the COST 1207 project in El Arenosillo, Spain,
74 six Brewers out of 18 differed by more than 10% from the reference, when using the
75 calibration provided by the operator
76 (http://www.pmodwrc.ch/wcc_uv/wcc_uv.php?topic=qasume_audit). The differences are most likely due
77 to slightly different corrections (for, e.g, temperature dependence and angular response) and
78 processing procedures. Small variations in a number of corrections and procedures may result
79 in large differences in the outcome.

80
81 For the comparability of the Brewer data from around the world, it is necessary to carefully
82 document the traceability of the calibration and the processing chain of the data. Rigorous
83 documentation should be a part of the routine QC/QA procedures at each site. This allows
84 anyone to audit all the steps taken prior to the delivery of the data, and allows making changes
85 in post-processing with no need to start everything from the beginning.

86
87 This study examines and demonstrates the steps that are involved in processing different
88 levels of solar spectral UV irradiance data produced by the Brewer spectrophotometers in
89 possession of the Finnish Meteorological Institute. Due to economical reasons, the Brewer
90 measurements at Jokioinen were terminated in November 2015. Since then, the Brewer #107
91 has been relocated and operated in Helsinki. Thus, this study also serves as a historical
92 description of the Jokioinen measurements, and as a platform for the development of the
93 procedures to be followed in Helsinki and Sodankylä in the future. A detailed description is

94 given on the process flow from the Level 0 data (raw counts) to the Level 2 data (quality
95 assured spectral UV irradiances and UV products). In a companion paper (Mäkelä et al. 2016)
96 we describe in detail how the final time series of the responsivity of a Brewer
97 spectrophotometer is derived. In another companion paper (Heikkilä et al. 2016) we describe
98 how the quality indicators provided by the European UV database (EUVDB) may be used for
99 quality assurance of Level 2 data.

100

101 2 Description of the stations

102 The Finnish Meteorological Institute (FMI) has been operating Brewer spectrophotometers at
103 two sites: Sodankylä and Jokioinen. In the following sections, brief descriptions on the
104 characteristics of the sites and their Brewer spectrophotometers are given.

105 2.1 The Sodankylä station (Brewer #037 and #214)

106 The Arctic Research Centre of the FMI is located at 67.37°N, 27.63°E, at 179 m of altitude
107 above sea level, in Sodankylä. It has maintained Brewer Mark II single monochromator
108 spectrophotometer since 1990. The nearby surroundings comprises of pine forests. In the
109 southwest flows the river Kitinen. The area is surrounded by a large peatland area in the east.
110 The terrain is snow covered typically from October to late April. The sun is just below the
111 horizon from mid-December to mid-January. Temperatures range from -40 °C in winter to
112 +30 °C in summer. The station is described in more detail in, e.g., Lakkala et al 2003.

113 Two Brewers are currently operated at this site: #037 and #214. They are located on the roof
114 of the sounding station (see Fig. 1). Brewer #037 has a single monochromator and a
115 wavelength range of 290-325 nm. The FWHM of the slit function is 0.56 nm. The other
116 Brewer, #214, has been set up in 2012, to work in tandem with Brewer #037 and to cover the
117 longer wavelengths of the UV radiation. The wavelength range of Brewer #214 is 286.5-365
118 nm. The FWHM of the slit function of Brewer #214 is 0.55 nm.

119 2.2 The Jokioinen station (Brewer #107)

120 The meteorological observatory in Jokioinen is located at 60.82°N, 23.50°E, at 107 m of
121 altitude above sea level, in a rural area surrounded by fields and mainly coniferous forests.
122 The ground is covered by snow most of the time during December-March. Temperatures can
123 range from -20 °C to +30 °C.

124 FMI acquired and set up Brewer #107 in Jokioinen in 1995. Since November 2015, it has
125 been relocated in Helsinki. At Jokioinen, Brewer #107 was located on the roof of the
126 sounding station (see Fig. 2). Brewer #107 is of type Mark III with a double monochromator.
127 The original wavelength range of the instrument was 286.5-363 nm. In April 1997, changes in
128 the optics were made. Since then, the wavelength range has been 286.5-365 nm. The FWHM
129 of the slit function is 0.59 nm.

130

131 3 Data flow

132

133 A schematic presentation on the data flow from the raw measurements performed by the FMI
134 Brewer spectrophotometers to the final UV products is given in Fig. 3. Three levels of data
135 are produced in the process. Level 0 data results from data acquisition, Level 1 data from the
136 near real time (NRT) processing of Level 0 data and the following quality control (QC), and
137 Level 2 data from post-processing of Level 0 data and the following quality assurance (QA).

138 To demonstrate the different phases in the data flow and the way the data is transformed as
139 they go through the whole chain from the raw counts to final products, we use one selected
140 case spectrum measured by Brewer #107 in Jokioinen on the 20th of May 2007. Information
141 related to the case scan is given in Table 1.

142

143 3.1 UV data acquisition – Level 0 data

144

145 The measurements performed by Brewer spectrophotometers are controlled through their own
146 operating computers. In UV irradiance measurements, the quartz dome viewport is used to
147 collect the photons from the hemisphere above. The wavelength range from the shortest
148 wavelength to the upper limit of the range is scanned in ascending direction with 0.5 nm
149 increments. The monochromator is used to select the photons on one narrow wavelength band
150 at a time and passed on to the photomultiplier tube (PMT) acting as a detector. The PMT
151 makes observations in cycles: four cycles at each wavelength below 300 nm and two cycles at
152 wavelengths above 300 nm. The raw counts are stored into files onto the operating computer

153 in units counts/cycle. These data are the Level 0 data produced by the Brewer
154 spectrophotometer. An example on the Level 0 data raw counts is shown in Fig. 4.

155

156 The operating computers of the Brewers work autonomously, making measurements based on
157 schedules predefined by the operator. Over the years, a variety of schedules has been used,
158 and the schedules for the two sites have supported slightly different research targets at
159 different times. For example, at Jokioinen, more frequent measurements have been made near
160 sunrise and sunset, and at constant air masses, as well as at time of the smallest solar zenith
161 angle (SZA). At Sodankylä, measurements have generally been spread out more evenly
162 throughout the day. At both sites, there has been a measurement at least every half hour
163 during daylight and a measurement at midday. The number of daily UV scans at Jokioinen,
164 for example, range from about 8 in winter to more than 30 in summer.

165

166 Other data acquired at the measurement sites and relevant to the UV data processing include
167 visibility and total ozone column. Automatic weather stations (AWS) produces data on
168 visibility every ten minutes. Total ozone column measurements are performed by the Brewer
169 spectrophotometers themselves. The data acquired on visibility and total ozone column are
170 used in the cosine correction procedure applied to the measured UV irradiance spectra
171 (Lakkala et al. 2008). As auxiliary measurements, a broadband UVB radiometer (Solar Light
172 SL501A) and a pyranometer (Kipp & Zonen CM11) measuring global radiation (305–2800
173 nm), are synchronized to the Brewer measurements. These measurements are used in offline
174 quality assurance (QA) procedures to identify erroneous measurements, and to obtain
175 information on changes in the cloud cover (Lakkala et al. 2008).

176

177 The Level 0 data on solar spectral UV irradiance are transferred from the Brewer's operating
178 computer to a central UNIX server. The raw data are further transferred to another server for
179 the quality control (QC) monitoring the proper functioning of the instrument. The system
180 taking care of the QC is called IDEAS (Integrated Data Evaluation & Analysis System) and is
181 briefly described in the following section. In addition, the raw data are uploaded every 20
182 minutes to the European Brewer Network Database (EUBREWNET,

183 <http://rbcce.aemet.es/eubrewnet>). The EUBREWNET is established within the COST 1207 project
184 of the European Union, as a joint effort of the international Brewer community.

185

186 **3.2 IDEAS – a quality control tool**

187

188 In 2015, new software was introduced to facilitate both immediate and long-term quality
189 control of data, and to improve the potential of the Brewers to work as real-time operational
190 devices. IDEAS (Integrated Data Evaluation & Analysis System, supplied by Full Spectrum
191 Science Inc.) is a tool using the Level 0 data directly for checking that the Brewer is
192 functioning correctly. The Brewer itself makes several check measurements during the day.
193 The measured parameters are used to monitor the stability of the instrument.

194

195 Every measurement of the Brewer and every process of the operating software is recorded in
196 appropriate data files. These are stored as so-called B-files, and updated to the server in which
197 IDEAS is running. IDEAS is integrated with the real-time 7/24 operational control system of
198 FMI. Automatic warnings of malfunctions of the Brewer are generated within 5-10 minutes
199 and sent to the 24/7 control center. If the personnel there are unable to solve the problems,
200 stand-by Brewer specialists can be alerted by text messages when needed.

201

202 The IDEAS software analyzes all records collected by the Brewer spectrophotometer
203 operational program (see Savastiouk 2011 for a description of an earlier version of the
204 software). These records include observations, tests and user comments. For each record
205 type that is known to the system, the results are compared to nominal or reference values,
206 prioritized and reported to the main FMI observation database system with flags indicating
207 the state of the instrument. For record types that are not yet known to the system a report is
208 created so new record types can be added to the system analysis. All data is stored in a
209 database for easy access at any time.

210

211 The full list of currently analyzed records contains over a hundred items. Some of the most
212 relevant critical tests include the wavelength calibration internal mercury lamp test (HG), the

213 linearity, or dead time, test (DT) and the spectral sensitivity stability test on the internal
214 halogen lamp (SL). Failures in any of these tests suggest that other data collected at the same
215 time are questionable and need attention from the operator and/or a scientist. In such cases,
216 an immediate alert is sent to the people responsible for the instrument to address these
217 failures.

218

219 Both real-time and historical reporting tool is also part of IDEAS. The tool generates a
220 collection of prioritized tables and plots that include test results as well as ozone data
221 comparison with satellites as an informational aid for the user to evaluate the data quality.

222

223 **3.3 UV data processing**

224

225 For the processing of solar spectral UV irradiance data produced by the Brewer
226 spectrophotometers, two separate lines have been established. The first is used for online
227 near-real-time (NRT) processing, using the Level 0 data and auxiliary data as inputs to
228 produce Level 1 data. The second is employed in offline post-processing, using the same data
229 as inputs but employing a re-evaluated time series of responsivity of the instrument,
230 correction for potential shifts in the wavelength scale, and final QA procedures. In both
231 processing schemes, knowledge on the responsivity of the instrument is needed for the
232 production of calibrated irradiance data.

233

234 **3.3.1 Calibration with Level 1 and Level 2 responsivities**

235

236 In order to convert the measured photon counts into irradiances, the response of the
237 instrument must be known. The response is determined using 1 kW lamp measurements
238 performed in an optical laboratory. The optical laboratory facilities at Sodankylä are described
239 in Lakkala et al. 2016. The optical laboratory facilities at Jokioinen are similar to those at
240 Sodankylä.

241

242 The calibration lamps are tungsten-filament incandescent halogen lamps of type DXW
243 operated in vertical beam direction. A primary standard is used to transfer the irradiance scale
244 from the National Standard Laboratory MIKES-Aalto. Using the measurements of the
245 Brewer, the scale is transferred to working standards used for the calibration of the Brewer
246 every six weeks. Usually at least three lamps are measured during one calibration event,
247 which enables the detection of potential drifts in the lamps (Webb et al. 1998, Lakkala et al.
248 2008). The obtained responsivity is used in the online (NRT) processing and production of
249 Level 1 data until the next calibration. Garane et al. 2006 have calculated that the typical
250 uncertainty of the UV irradiances due to the absolute calibration uncertainty is 1.4%.

251
252 The determination of the Level 2 responsivity with daily values is briefly described in
253 Lakkala et al. (2008) and is covered in more detail in the companion paper (Mäkelä et al.
254 2016). In brief, calibration lamp measurements are analysed back in time on a longer time
255 span and measurements of all lamps are incorporated in cross-checking and comparison, to
256 separate the long-term drifts of the lamps from transient exceptions in their output, and to
257 confirm the selection of the core “trusted” lamps that serve as the basis for the final Level 2
258 calibration. This analysis results in a response time series with daily responsivities over the
259 time period investigated. The time series is formed by linear interpolation between the
260 moments of lamp measurements and smoothed by using a moving average.

261
262 Examples on the responsivities are shown in Fig. 4. The Level 1 and Level 2 responsivities
263 plotted in the figure have been applied in the online (NRT) and offline post-processing of the
264 raw counts recorded by Brewer #107 on 20 May 2007 (the case scan described in Table 1).
265 The interpolated and smoothed time series of Level 2 responses at 305 nm for Brewer #037 is
266 shown in Fig. 5a to demonstrate the temporal development of the response of the instrument.
267 In Fig. 5b, the stepwise time series of Level 1 and the interpolated and smoothed time series
268 of Level 2 responsivities at 305 nm for Brewer #107 are also shown partly overlapping to
269 illustrate the difference between the responsivities used in Level 1 and Level 2 processing.

270
271 Between the 1 kW lamp calibrations performed in the laboratory, the stability of the Brewer is
272 monitored on-site every three weeks using external 50 W lamp measurements. In those
273 measurements, at least three lamps are used as well. If a change in the spectral response of the
274 Brewer is detected, the Brewer is moved to the laboratory for absolute calibration.

275

276 3.3.2 Processing algorithms

277

278 Both the online and offline data processing are done using custom-made software written in
279 Perl (Practical Extraction and Report Language) and supported by shell scripts. The original
280 software provided by the instrument manufacturer and the algorithms therein has served as a
281 basis for the software development. The algorithms have been described in detail by Lakkala
282 et al. (2008). Only the key points are therefore summarized in the following.

283

284 The Brewers perform the scans by starting from the shortest and ending at the longest
285 wavelength of the wavelength range of the instrument. The Brewer software also returns the
286 dark current, F1, for each scan. The total scanning time is about 4 minutes for Brewer #037
287 and 5 minutes for Brewers #107 and #214 due to the larger wavelength range. The
288 occasionally occurring noise spikes are first removed based on the method of Meinander et al.
289 (2003). Following SCI-TEC Instruments Inc. (1999), the raw counts F (in units counts/cycle)
290 recorded by the Brewer are next converted to count rates C (counts/second) by subtracting the
291 dark counts F1 and taking into account the integration time IT and the number of cycles CY
292 of the slit exposure. The integration time is pre-defined as 0.1147 s. For each wavelength i,
293 where $i=290-325(365)\text{nm}$, the dark current corrected count rate is calculated as

294

$$295 \quad C_i = \frac{2(F_i - F1)}{CY * IT} \quad (1)$$

296

297 The dead time DT of the photomultiplier is measured daily, and taken into account using
298 Poisson statistics. Nine iteration rounds (n=0-8) are next used to correct the count rate for the
299 dead time:

$$300 \quad C_i(n) = C_i(0) * e^{C_i(n) * DT} \quad (2)$$

301

302 The stray light is calculated as the average of all counts below 292 nm (#107 and #214) or
303 293 nm (#037). The count rate corrected for stray light is simply

304
$$C_i = C_i - \frac{1}{N} \sum_{i < 292(3)} C_i \quad (3)$$

305 where N is the number of wavelengths below the stray light limit. Since #107 and #214 have a
306 double monochromator, the stray light counts are small. The Sodankylä Brewer #037 has a
307 single monochromator, and the stray light counts are larger (Bais et al. 1996).

308

309 The count rates are then converted to irradiances by dividing the count rate C_i with the
310 spectral response R_i of the instrument:

311
$$I_i = \frac{C_i}{R_i} \quad (4)$$

312 Temperature and cosine corrections are then made to the spectral irradiances. A linear
313 dependence for temperature is assumed:

314
$$I_i = I_i * (1 + CT_i * \Delta T), \quad (5)$$

315 where CT_i is the temperature correction coefficient and ΔT is the temperature difference
316 between the measurement temperature and the reference temperature.

317

318 The cosine correction is based on radiative transfer calculations (Lakkala et al. 2008). As both
319 the direct and the diffuse component of the global radiation affect the cosine correction factor,
320 we need to estimate the ratio of the direct and diffuse components of the actual global
321 radiation. This is done using the freely available libRadtran package (Mayer and Kylling
322 2005). The only unknown input parameter to the model is the cloud optical depth, which is
323 estimated using a predefined lookup table. The inputs to the lookup table include the total
324 ozone (measured with the Brewer), the irradiance given by the Brewer, the solar zenith angle,
325 the visibility (given by the AWS), and the albedo. Currently, a fixed albedo value of 0.03 is
326 used for both stations, corresponding to the average summer albedo.

327

328 **3.3.3 Online processing - Level 1 data**

329

330 The components of the online NRT processing scheme are described in Fig. 6. Processing of
331 all the data from both stations is done at a single central Unix server. The raw (Level 0) data

332 is transferred from the Brewer operational computer to the central server every five minutes.
333 From the central computer the data is further transferred to the IDEAS server, where the QC
334 of the data is done. Every five to ten minutes status messages are sent to the observation
335 control server, from which the critical status messages are transferred to the 24/7 control of
336 FMI. Level 0 data is also transferred to the European Brewer database, EUBREWNET every
337 20 minutes. As soon as a new UV measurement has been detected by the central server, the
338 data processing of Level 1 data starts and visibility is downloaded from the climate database
339 of FMI, and used for the cosine correction of the data. The Level 1 data is then used to
340 calculate products which are transferred to a server with a web interface and the climate data
341 base of FMI. The reference data from the broadband SL501A radiometer and the global
342 radiation pyranometer CM11 are downloaded each time a spectral measurement is made. An
343 example on the Level 1 irradiance spectrum is given in Fig. 7.

344

345 The Level 1 data is spectral UV irradiance processed in near real time. Several corrections are
346 routinely made in the Level 1 processing, and several data products are derived. In addition to
347 the near-real-time measurements, every morning the data from the previous day is
348 reprocessed. This is mainly done to avoid data gaps in Level 1 data. Such gaps might for
349 example occur if there has been a malfunction of the central server, in which case the Level 1
350 spectra have not been calculated even though the Level 0 data from the Brewer are still
351 available.

352

353

354 **3.3.4 Offline processing - Level 2 data**

355

356 Level 1 spectral irradiance data is produced using the Level 1 responsivity updated regularly
357 on the basis of the calibration lamp measurements. To produce Level 2 data, calibration lamp
358 measurements are analysed back in time to produce response time series as described in
359 Section 3.3.1. By this procedure, the potential drifts in the calibration lamps, possibly
360 affecting the Level 1 responsivity, are eliminated. This correction can be calculated once a
361 year, when the primary calibration lamps are recalibrated. The Level 2 data is fully
362 homogenized and quality assured. The SHICRIVM algorithm (Slaper et al 1995) is used to
363 correct for wavelength shifts. As a quality control tool, the dose rates are compared with

364 reconstructed UV, model calculations of clear sky UV and global radiation, global radiation
365 and broad band UV data in order to distinguish erroneous measurements (Lakkala et al.
366 2008). In addition each spectrum is checked by eye and bad measurements are excluded. An
367 example on the Level 2 irradiance spectrum is given in Fig. 7.

368
369 Once a year the Level 2 data is uploaded to the database EUVDB (<http://uv.fmi.fi/uvdb/>). Figure
370 8 shows the number of UV spectra submitted to the EUVDB measured by the Brewer #037
371 (Fig. 8a) and Brewer #107 (Fig. 8b) during 1990-2014 and 1995-2014, respectively. The
372 submission of Brewer #037 spectra for 2011 will be upgraded in near future, as there exists
373 more than 5700 spectra for that year, but only some of them have been submitted to the
374 database at the time of writing the manuscript. The data from Sodankylä is also uploaded to
375 the database of the FMI-Arctic Research Centre (<http://litdb.fmi.fi/>).

376

377 **3.3.5 Products**

378

379 Several products are derived from the Level 1 and Level 2 spectral data. These include dose
380 rates (in $W\ m^{-2}$) and daily doses (in $J\ m^{-2}$) as unweighted and weighted by selected action
381 spectra (Table 2 and Fig. 9). Let us denote an arbitrary action spectrum by $s(\lambda)$. The dose rate
382 \dot{D} is calculated from the irradiance spectrum, multiplied by the action spectrum, by numerical
383 integration over the appropriate wavelength range (UV, UVA or UVB):

$$384 \quad \dot{D} = \int_{\lambda_{min}}^{\lambda_{max}} s(\lambda) \cdot E(\lambda) d\lambda. \quad (6)$$

385 For the unweighted quantities, $s(\lambda)$ is equal to unity. The daily dose is further derived from
386 the dose rate by numerical integration over the day:

$$387 \quad D = \int_{t_{min}}^{t_{max}} \dot{D} dt. \quad (7)$$

388 The derivation of the products is illustrated in Fig. 10. The action spectra currently in routine
389 use are listed in Table 2. UV index is additionally derived by multiplying the CIE erythemally
390 weighted UV dose rate by 40 (WMO 1997).

391

392 For the calculation of the dose rates requiring integration beyond the upper wavelength limit

393 of the Brewer, the measured spectra are extended using a pre-defined reference UVA
394 spectrum (Fig. 10). The extension is adjusted onto the level of the measured spectrum by
395 linear conversion. The ratio of the measured irradiance to the reference irradiance at selected
396 wavelength is used as a scaling factor. For Brewer #037, the wavelength of 324 nm, and for
397 Brewer #107 and #214, the wavelength of 361 nm is used as a point of adjustment. All action
398 spectra in routine processing (Fig. 9) approach zero towards the longer UVA wavelengths.
399 This means that the uncertainty caused by the artificial UVA extension to the computed dose
400 rate is of the order of 10^{-3} . For the unweighted UV and UVA dose rates, our investigation
401 based on a radiative transfer model simulation suggest uncertainties as high as approx. 2 %
402 caused by the constant scaled UVA extension. This finding is in line with the result obtained
403 by Fioletov et al. 2004.

404

405 All products are stored in the central server, and graphs and statistical tables are updated each
406 time a new measurement is processed. The data are transferred to a server which has a web
407 interface for internal use at FMI. This information is used at FMI for operational and research
408 purposes. In addition, the Level 1 UV index is transferred NRT to the FMI climate database.

409

410

411 4 **Conclusions**

412 Production of high quality data on solar spectral UV at the Earth's surface requires a complex
413 set of operation routines, processing algorithms, and QC/QA procedures. We have described
414 and demonstrated the methods used in the acquisition and processing of solar spectral UV
415 irradiance data measured by Brewer spectrophotometers in Finland. As a result, we have
416 produced a comprehensive view on the multi-phase data flow from the collection of single
417 UV photons to calculation of final products of UV irradiance. This is expected to facilitate
418 identification and isolation of specific targets of development in the procedures used in
419 different phases of the data flow. The first targets of this kind appear to be the implementation
420 of solar zenith angle and total column ozone dependent UVA extension to the measured
421 spectra in the calculation of the different UV dose rates. Another emerging idea has been
422 development of a comprehensive data and metadata management system to support the data
423 flow as a whole and to ensure its continuance.

424

425 Due to economical reasons, solar spectral UV measurements at the Jokioinen observatory
426 were terminated in November 2015, and the Brewer #107 was transferred to and set up for
427 measurements in Helsinki, the capital city of Finland. The twenty-year time series of solar
428 spectral UV irradiance obtained in Jokioinen is in the process of being analysed. The methods
429 and procedures reported in this study are still in continuous use in Helsinki and Sodankylä.
430 Brewer #037 is still operated in Sodankylä and continues to collect its over 25-year record of
431 solar spectral UV measurements. In Helsinki, a new time series has been initiated with
432 Brewer #107. Thorough understanding on the different phases in the data flow and further
433 development of the methods is a pre-requisite for the vitality of the monitoring programs also
434 in the future.

435

436 **Acknowledgements**

437 Professor Esko Kyrö is acknowledged for starting the Brewer measurements at Sodankylä.
438 We are grateful to the operators at Sodankylä and Jokioinen stations for daily maintenance
439 and for performing the calibrations of the Brewers. We thank the Brewer community within
440 the COST 1207 project for sharing expertise related to Brewer measurements.

441

442 **References**

443 Bais, A. F., Gardiner, B. G., Slaper, H., Blumthaler, M., Bernhard, G., McKenzie, R., Webb,
444 A. R., Seckmeyer, G., Kjeldstad, B., Koskela, T., Kirsch, P. J., Gröbner, J., Kerr, J. B.,
445 Kazadzis, S., Leszczynski, K., Wardle, D., Josefsson, W., Brogniez, C., Gillotay, D., Reinen,
446 H., Weihs, P., Svenoe, T., Eriksen, P., Kuik, F. and Redondas, A.: SUSPEN intercomparison
447 of ultraviolet spectroradiometers, *J. Geophys. Res.*, 106(D12), 12509–12525,
448 doi:10.1029/2000JD900561.

449 Bais, A., Zerefos, C. and McElroy, C.: Solar UVB measurements with the double- and single-
450 monochromator Brewer Ozone Spectrophotometers, *Geophys. Res. Lett.*, 23, 833–836, 1996.

451 Bernhard, G., Dahlback, A., Fioletov, V., Heikkilä, A., Johnsen, B., Koskela, T., Lakkala, K.
452 and Svendby, T.: High levels of ultraviolet radiation observed by ground-based instruments
453 below the 2011 Arctic ozone hole, *Atmos. Chem. Phys.*, 13, 10573–10590, www.atmos-chem-
454 phys.net/13/10573/2013/ doi:10.5194/acp-13-10573-2013, 2013.

455 Bernhard, G. and Seckmeyer, G.: Uncertainty of measurements of spectral solar UV
456 irradiance, *J. Geophys. Res.*,104, D12, 14,321-14,345, 1999.

457 Brewer, A. W.: A replacement for the Dobson spectrophotometer?. *Pure Appl. Geophys.*,
458 106-108, 919–927, 1973.

459 Caldwell, M. M., Camp, L.B., Warner, C. W. and Flint, S.D.:Action spectra and their key role
460 in assessing biological consequences of solar UV-B radiation change. In: Worrest, R.C. and
461 M.M. Caldwell (Eds.): Stratospheric ozone reduction, solar ultraviolet radiation and plant life.
462 Springer-Verlag, Berlin. 87-111, 1986.

463 Cappellani, F. and Bielli, A.: Correlation between SO₂ and NO₂ measured in an atmospheric
464 column by a Brewer spectrophotometer and at ground-level by photochemical techniques,
465 *Environmental Monitoring and Assessment*, Vol 35, 2, 77-84, 1995.

466 Cede, A., J. Herman, A. Richter, N. Krotkov, and Burrows, J.: Measurements of nitrogen
467 dioxide total column amounts using a Brewer double spectrophotometer in direct Sun mode,
468 *J. Geophys. Res.* 111, D05304, doi:10.1029/2005JD006585, 2006.

469 CIE (International Commission on Illumination), Action spectrum for the production of
470 previtamin D3 in human skin, *CIE*, 174, 2006.

471 Diémoz, H., Siani, A. M., Redondas, A., Savastiouk, V., McElroy, C. T., Navarro-Comas, M.,
472 and Hase, F.: Improved retrieval of nitrogen dioxide (NO₂) column densities by means of
473 MKIV Brewer spectrophotometers, *Atmos. Meas. Tech.*, 7, 4009-4022, doi:10.5194/amt-7-
474 4009-2014, 2014.

475 Eleftheratos K., Kazadzis, S., Zerefos, C. S., Tourpali, K., Meleti, C., Balis, D., Zyrichidou,
476 I., Lakkala, K., Feister, U., Koskela, T., Heikkilä, A. and Karhu, J. M.: Ozone and
477 Spectroradiometric UV Changes in the Past 20 Years over High Latitudes, *Atmosphere-*
478 *Ocean*, DOI: 10.1080/07055900.2014.919897, 2014.

479 Fioletov, V. E., M. G. Kimlin, N. Krotkov, L. J. B. McArthur, J. B. Kerr, D. I. Wardle, J. R.
480 Herman, R. Meltzer, T. W. Mathews, and J. Kaurola: UV index climatology over the United
481 States and Canada from ground-based and satellite estimates, *J. Geophys. Res.*, 109, D22308,
482 doi:10.1029/2004JD004820, 2004.

483 Garane, K., Bais, A. F., Kazadzis, S., Kazantzidis, A., and Meleti, C.: Monitoring of UV
484 spectral irradiance at Thessaloniki (1990–2005): data re-evaluation and quality control *Ann.*
485 *Geophys.*, 24, 3215–3228, 2006.

486 Grobner, J., R. Vergaz, V. E. Cachorro, D. V. Henriques, K. Lamb, A. Redondas, J. M.
487 Vilaplana, and Rembges, D.: Intercomparison of aerosol optical depth measurements in the
488 UVB using Brewer spectrophotometers and a Li-Cor spectrophotometer, *Geophys. Res. Lett.*
489 28, 1691-1694, 2001.

490 de Gruijl, F.R. and van der Leun, J. C.: Estimate of the wavelength dependency of ultraviolet
491 carcinogenesis in humans and its relevance to the risk assessment of a stratospheric ozone
492 depletion. *Health Phys.*, vol 67, 319-325, 1994.

493 Hassinen S., Tamminen, J., Tanskanen, A., Koskela, T., Karhu, J. M., Lakkala, K. and
494 Mälkki, A.: Description and Validation of the OMI Very Fast Delivery Products, *J. Geophys.*
495 *Res.*, 113, D16S35, doi:10.1029/2007JD008784, 2008.

496 Heikkilä, A., Methods for assessing degrading effects of UV radiation on materials, Finnish
497 Meteorological Institute Contributions, 111, ISBN 978-951-697-843-0, Unigrafia Oy,
498 Helsinki, 2014.

499 Heikkilä, A., Kaurola, J., Lakkala, K., Karhu, J.M., Kyrö, E., Koskela, T., Engelsen, O.
500 Slaper, H. and Seckmeyer, G., European 1 UV DataBase (EUVDB) as a repository and
501 quality analyzer for solar spectral UV irradiance monitored in Sodankylä, *Geosci. Instrum.*
502 *Method. Data Syst. Discuss.*, doi:10.5194/gi-2015-39, in review, 2016.

503 Josefsson, W., Koskela, T., Dahlback, A. and Eriksen, P.: Spectral sky measurements. In *The*
504 *Nordic intercomparison of ultraviolet and total ozone instruments at Izaña from 24 October to*
505 *5 November 1993. Final report, edited by Koskela, T., Finnish Meteorological Institute,*
506 *Meteorological publications No. 27, Helsinki, 73-80, 1994.*

507 Kazadzis, S., Bais, A., Kouremeti, N., Gerasopoulos, E., Garane K., Blumthaler, M.,
508 Schallhart, B. and Cede A.: Direct spectral measurements with a Brewer spectroradiometer:
509 Absolute calibration and aerosol optical depth retrieval, *Appl. Opt.*, 44(9), 1681 – 1690, 2005.

510 Kazantzidis, A., Bais, A., Zempila, M., Kazadzis, S., den Outer, P., Koskela, T. and Slaper, H.:
511 Calculations of the human Vitamin D exposure from UV spectral measurements at three
512 European stations, *Photochem. Photobiol. Sci.*, 2009, 8, 45-51, 2009.

513 Knudsen, B., Larsen, N., Mikkelsen, I., Morcette, J.-J., Braahten, G., Kyrö, E., Fast, H.,
514 Gernand, H., Kanzawa, H., Nakane, H., Dorokhov, V., Yushkov, V., Hanse, G., Gil, M. and
515 ShearmanR.: Ozone depletion in and below the Arctic vortex for 1997. *Geophys. Res. Lett.*,
516 25, 627-630, 1998.

517 Koskela, T., Johnsen, B., Bais, A., Josefsson, W. and Slaper, H., 1997: Spectral sky
518 measurements. In *Nordic intercomparison of ultraviolet and total ozone instruments at Izaña*
519 *October 1996. Final report*, edited by Kjeldstad, B., Johnsen, B. and Koskela, T., Finnish
520 Meteorological Institute, Meteorological publications No. 36, Helsinki, 109-148, 1997.

521 Lakkala, K., Kyrö, E. and Turunen, T: Spectral UV Measurements at Sodankylä during 1990–
522 2001, *Journal of Geophysical Research*, 108, D19, 4621, doi:10.1029/2002JD003300, 2003.

523 Lakkala, K., Arola, A., Heikkilä, A., Kaurola, J., Koskela, T., Kyrö, E., Lindfors, A. V.,
524 Meinander, O., Tanskanen, A., Gröbner, J. and Hülsen, G.: Quality assurance of the Brewer
525 spectral UV measurements in Finland, *Atmos. Chem. Phys.*, 8, 3369–3383, 2008.

526 Lakkala, K., Suokanerva, H., Karhu, J.M., Aarva, A., Poikonen, A., Karppinen, T., Ahponen,
527 M., Hannula, H.-R., Kontu, A. and Kyrö, E.: Optical laboratory facilities at the Finnish
528 Meteorological Institute – Arctic Research Centre. *Geosci. Instrum. Method. Data Syst.*
529 *Discuss.*, doi:10.5194/gi-2015-43, 2016.

530 Lappalainen N., Huttunen, S., Suokanerva, H. and Lakkala, K.: Seasonal acclimation of the
531 moss *Polytrichum juniperinum Hedw.* to natural and enhanced ultraviolet radiation, *Environ*
532 *Pollut.*, 158, 3, 891-900, 2010.

533 Manney G., Santee M., Rex M., Livesey N., Pitts M., Veefkind P., Nash E., Wohltmann I.,
534 Lehmann R., Froidevaux L., Poole L., Schoeberl M., Haffner D., Davies J., Dorokhov V.,
535 Gernandt H., Johnson B., Kivi R., Kyrö E., Larsen N., Levelt P., Makshtas A., McElroy T.,
536 Nakajima H., Parrondo M., Tarasick D., von der Gathen P., Walker K. and Zinoviev N.:
537 Unprecedented Arctic ozone loss in 2011, *Nature*, 478, 469–475, doi:10.1038/nature10556,
538 2011.

539 Marengo, F., A. di Sarra, and De Luisi, J.: Methodology for determining aerosol optical depth
540 from Brewer 300-320-nm ozone measurements, *Appl. Opt.*, 41, 1805-1814, 2002.

541 Martz F., Turunen, M., Julkunen-Tiitto, R., Lakkala, K. and Sutinen, M.-L.: Effect of the
542 temperature and the exclusion of UVB radiation on the phenolics and iridoids in *Menyanthes*

543 *trifoliata* L. leaves in the subarctic, *Environmental Pollution*, 157, 3471-3478,
544 10.1016/j.envpol.2009.06.022, 2009.

545 Mayer, B. and A. Kylling, Technical note: The libRadtran software package for radiative
546 transfer calculations – description and examples of use, *Atmos. Chem. Phys.*, 5, 1855–1877,
547 2005

548 McKinlay, A. F. and Diffey, B. L.: A reference action spectrum for ultraviolet induced
549 erythema in human skin. CIE Research Note, CIE-Journal, Vol. 6, No.1(17-22), 1987.

550 Meinander, O., Josefsson, W., Kaurola, J., Koskela, T. and Lakkala, K.: Spike detection and
551 correction in Brewer spectroradiometer ultraviolet spectra, *Opt. Eng.* 42, 6, 1812–1819, 2003.

552 Mitchell, B.G., "Action Spectra for ultraviolet photoinhibition of Antarctic phytoplankton and
553 a model of spectral diffuse attenuation coefficients", In *Response of Marine Phytoplankton to*
554 *Natural Variations in UV-B Flux*, (Edited by G. Mitchell, I. Sobolev and O. Holm-Hansen),
555 Proc. of Workshop, Scripps Institution of Oceanography, La Jolla, CA, April 5, 1990

556 Munakata, N., Kazadzis, S., Bais, A.F., Hieda, K., Rontó, G., Rettberg, P., and Horneck, G.:
557 Comparisons of spore dosimetry and spectral photometry of solar-UV radiation at four sites in
558 Japan and Europe. *Photochem. Photobiol.* 72, 739-745, 2000.

559 Mäkelä, J.S., Lakkala, K., Meinander, O., Kaurola, J. Koskela, T., Karhu, J.M., Karppinen,
560 T., Kyrö, E., Leeuw, G. and Heikkilä, A.: In search of traceability: Two decades of calibrated
561 Brewer UV measurements in Sodankylä and Jokioinen, *Geosci. Instrum. Method. Data Syst.*
562 *Discuss.*, doi:10.5194/gi-2015-40, in review, 2016.

563 Savastiouk, V: A database implementation of data analysis and quality control for the Brewer,
564 presentation at the 13th Brewer User Workshop. Beijing, China September 12-16, 2011
565 ([https://www.wmo.int/pages/prog/arep/gaw/documents/13th_Brewer_d2_Savastiouk-](https://www.wmo.int/pages/prog/arep/gaw/documents/13th_Brewer_d2_Savastiouk-Database.pdf)
566 [Database.pdf](https://www.wmo.int/pages/prog/arep/gaw/documents/13th_Brewer_d2_Savastiouk-Database.pdf))

567 SCI-TEC Instruments Inc.: Brewer MKII Spectrophotometer, operator's manual, Saskatoon,
568 Sask. , Canada, 1999.

569 Setlow, R.B.: The wavelengths in sunlight effective in producing skin cancer: A theoretical
570 analysis. *Proc.Natl.Acad.Sci., U.S.A.* 71:3363-3366, 1974.

571 Slaper, H., Reinen, A. J., Blumthaler, M., Huber, M., and Kuik, F.: Comparing ground-level
572 spectrally resolved solar UV measurements using various instruments: A technique resolving
573 effects of wavelength shift and slit width, *Geophys. Res. Lett.*, 22, 2721–2724, 1995.

574 Webb, A., Gardiner, B., Leszczynski, K., Mohnen, V. A., Johnston, P., Harrison, N. and
575 Bigelow, D.: Quality Assurance in Monitoring Solar Ultraviolet Radiation: the State of the
576 Art. World Meteorological Organization (WMO), Global Atmosphere Watch Report, 146,
577 2003.

578 Webb, A. R., Gardiner, B. G., Martin, T. J., Leszczynski, K., Metzdoeff, J., Mohnen, V. A. and
579 B. Forgan, Guidelines for site quality control of UV monitoring, Rep.Ser. 126, *Environ.*
580 *Pollut. Monit. And Res. Programme*, World Meteorol. Organ., Geneva, 1998.

581 WMO (World Meteorological Organization), Report of the WMO-WHO Meeting of Experts
582 on UVB Measurements, Data Quality and Standardisation of UV indices, Global Atmosphere
583 Watch Report No. 95, 1997.

584

585

586

587

588

589

590

591

592

593



594

595 Figure 1. View of Sodankylä observatory. The Brewer is located on the roof in the enclosure
596 on the left.

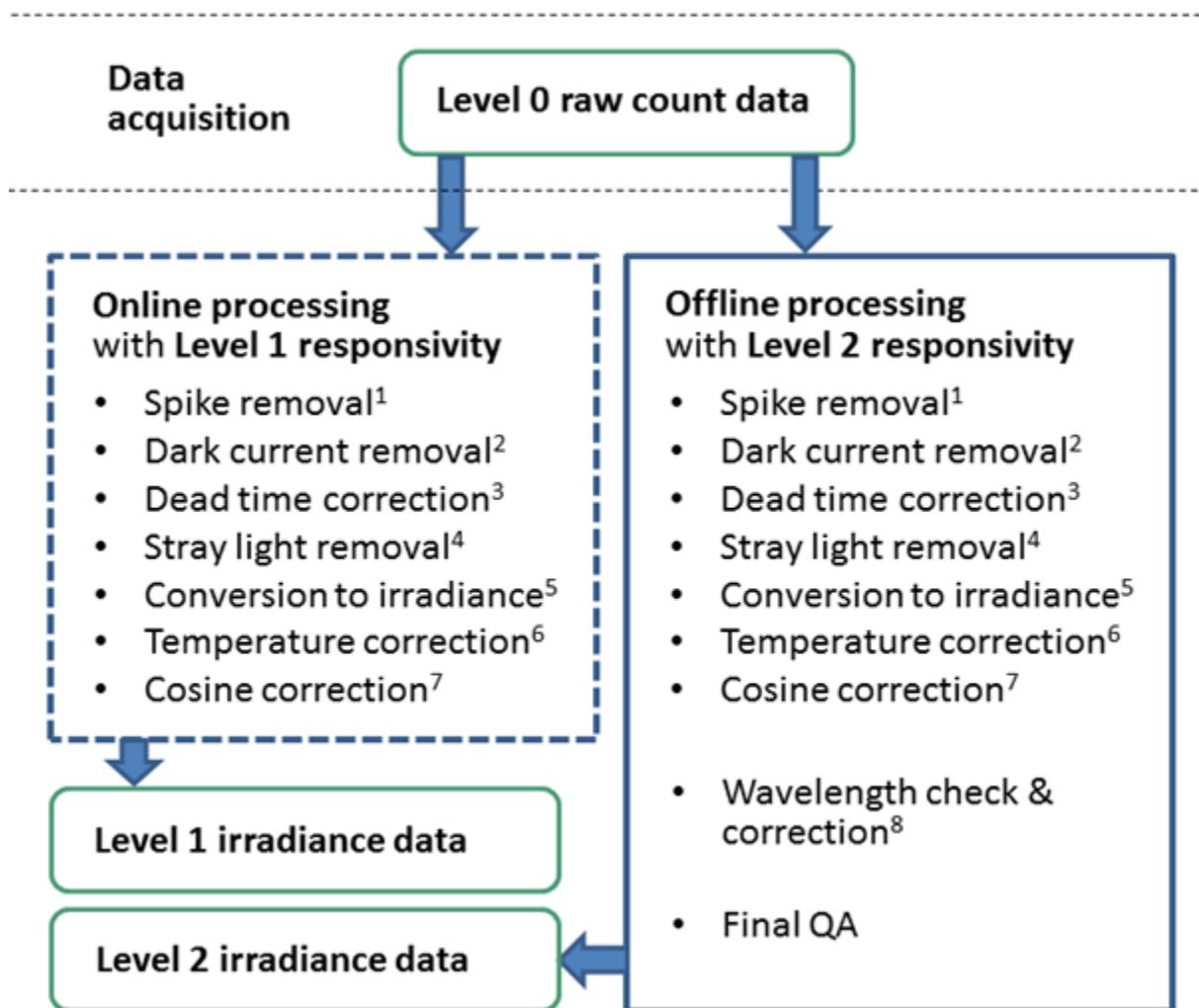
597



598

599

600 Figure 2. View of Jokioinen observatory. The Brewer is located on the roof in the enclosure
601 on the left.

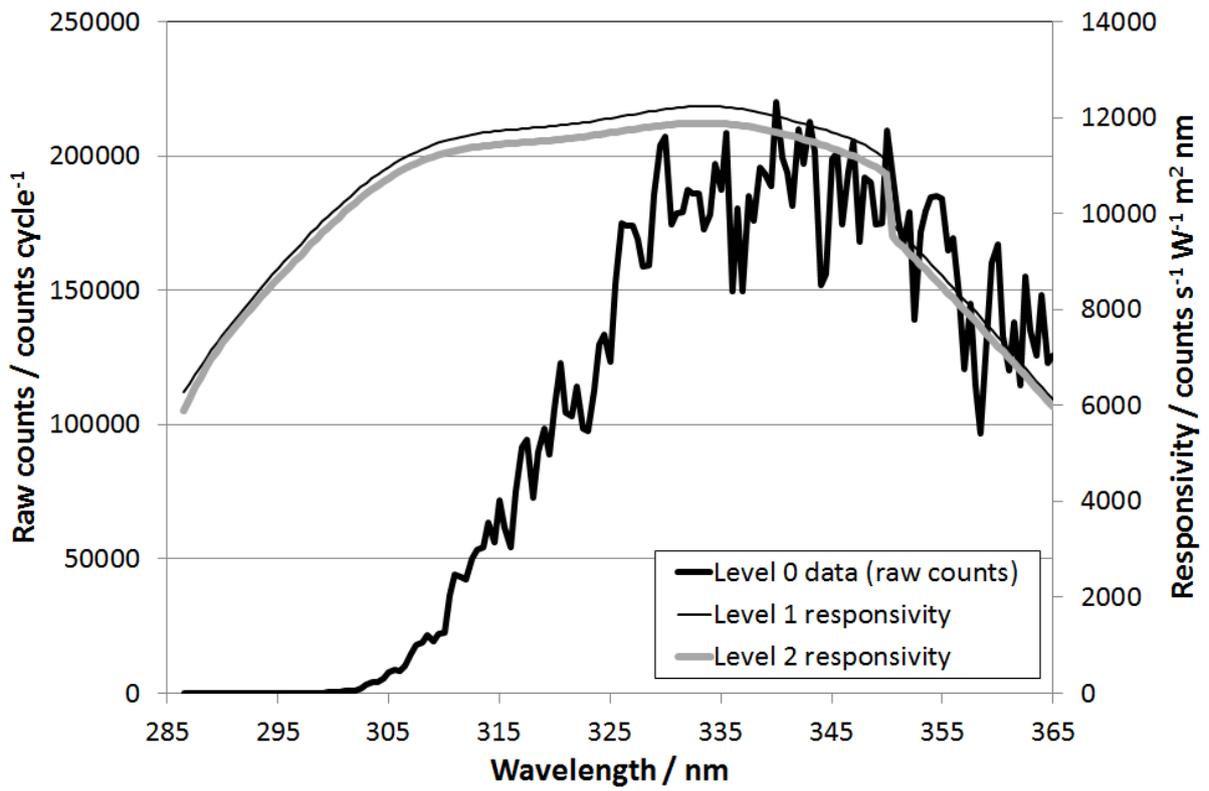


602

603 Figure 3. Schematic presentation on the data flow of Brewer UV irradiance data from the
 604 Level 0 raw counts to Level 1 and Level 2 irradiance data.

605 ¹Meinander et al. 2003; ²Eq. 1; ³Eq. 2; ⁴Eq. 3; ⁵Eq. 4; ⁶Eq. 4; ⁷Lakkala et al. 2008; ⁸Slaper et
 606 al. 1995.

607

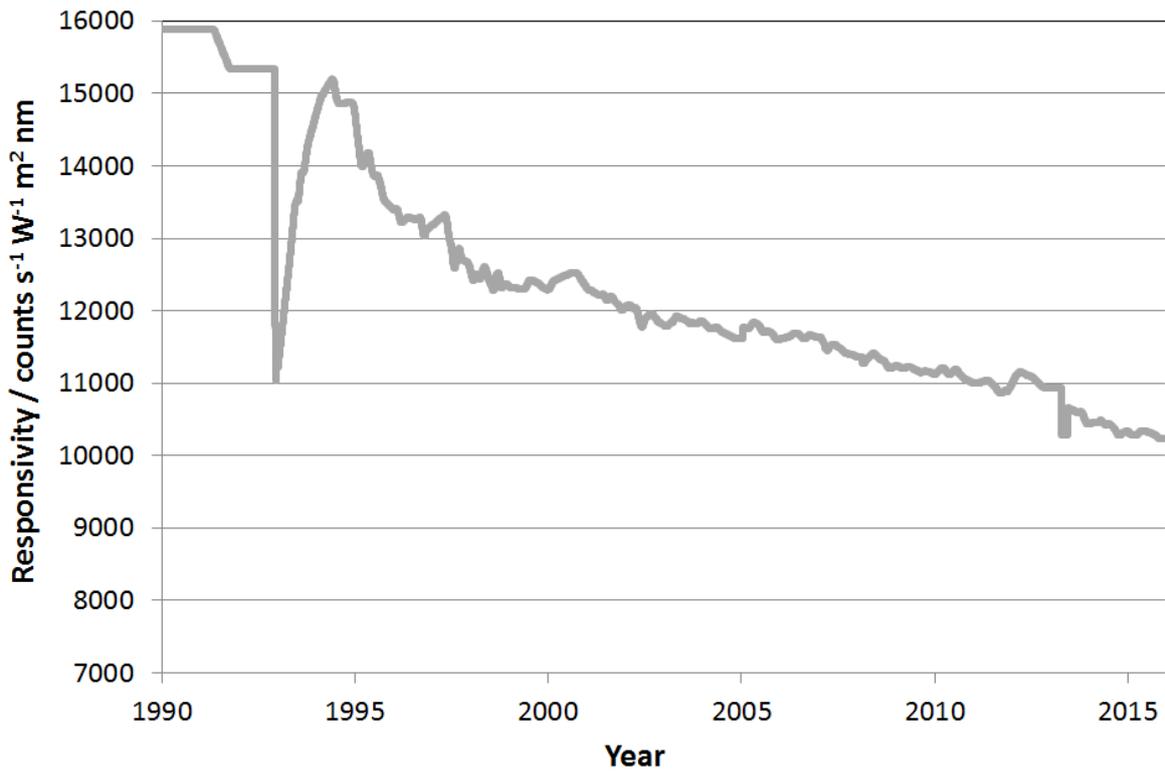


608

609

610 Figure 4. Example of Level 0 raw count data recorded by the Brewer #107 on 20 May 2007 at
 611 13:20:07–13:25:18 UTC (on primary vertical axis) and Level 1 / Level 2 responsivities of the
 612 instrument for the same day (on secondary vertical axis).

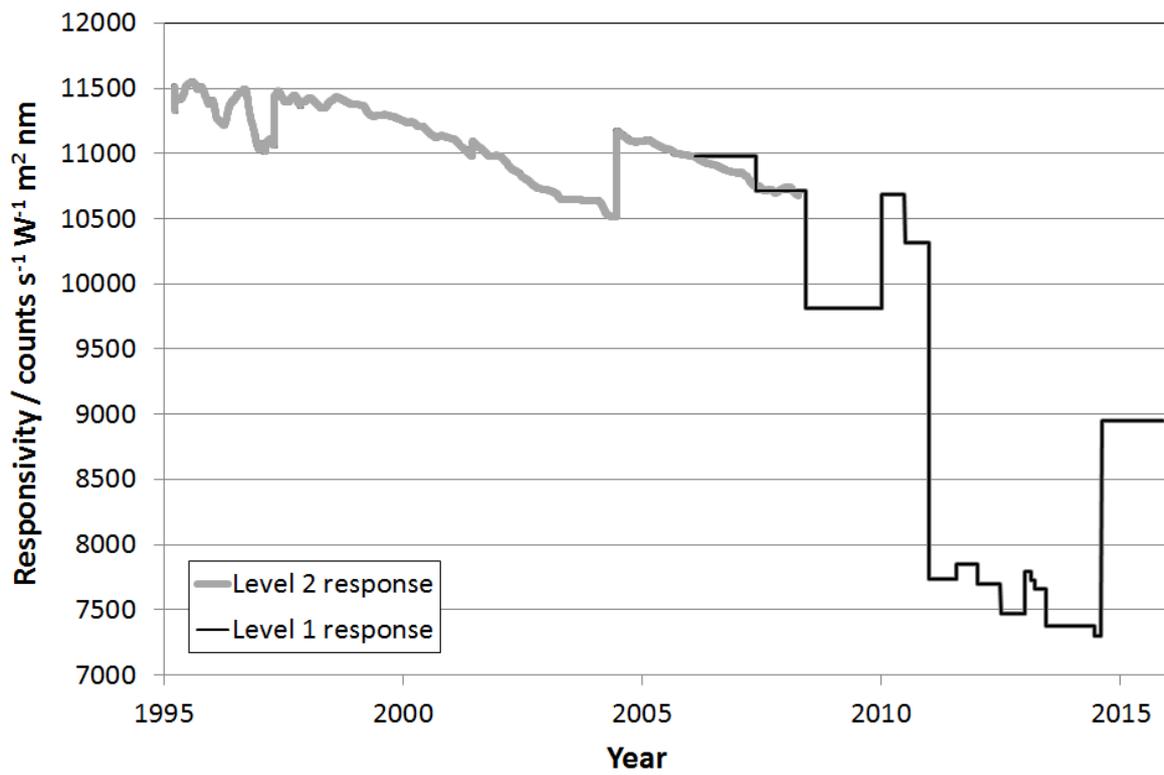
613



614

615 Figure 5a. Time series of Level 2 responsivity of Brewer #037 at 305 nm.

616



617

618 Figure 5b. Partly overlapping time series of Level 1 and Level 2 responsivities of Brewer

619 #107 at 305 nm.

620

621

622

623

624

625

626

627

628

629

630

631

632

633

634

635

636

637

638

639

640

641

642

643

644

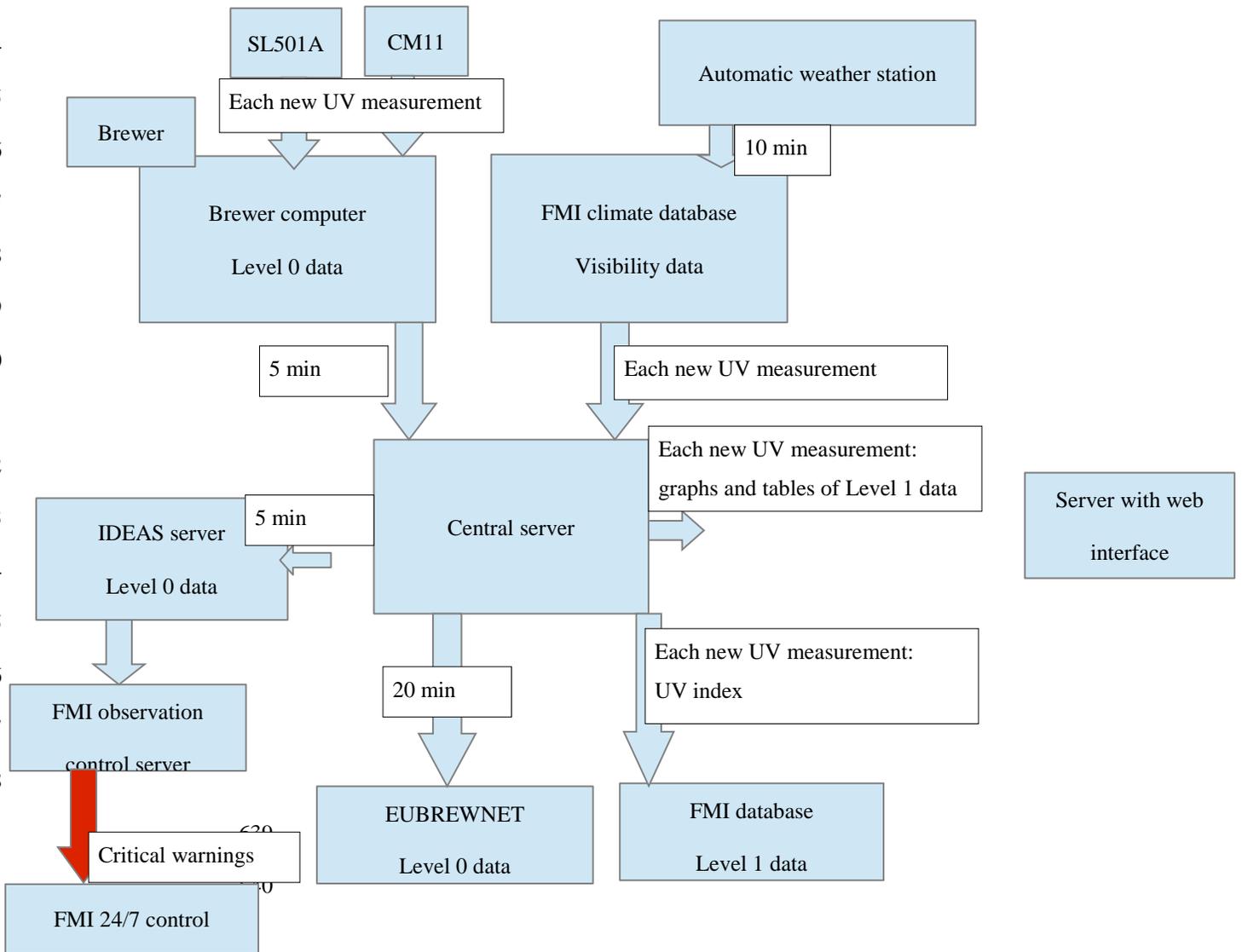
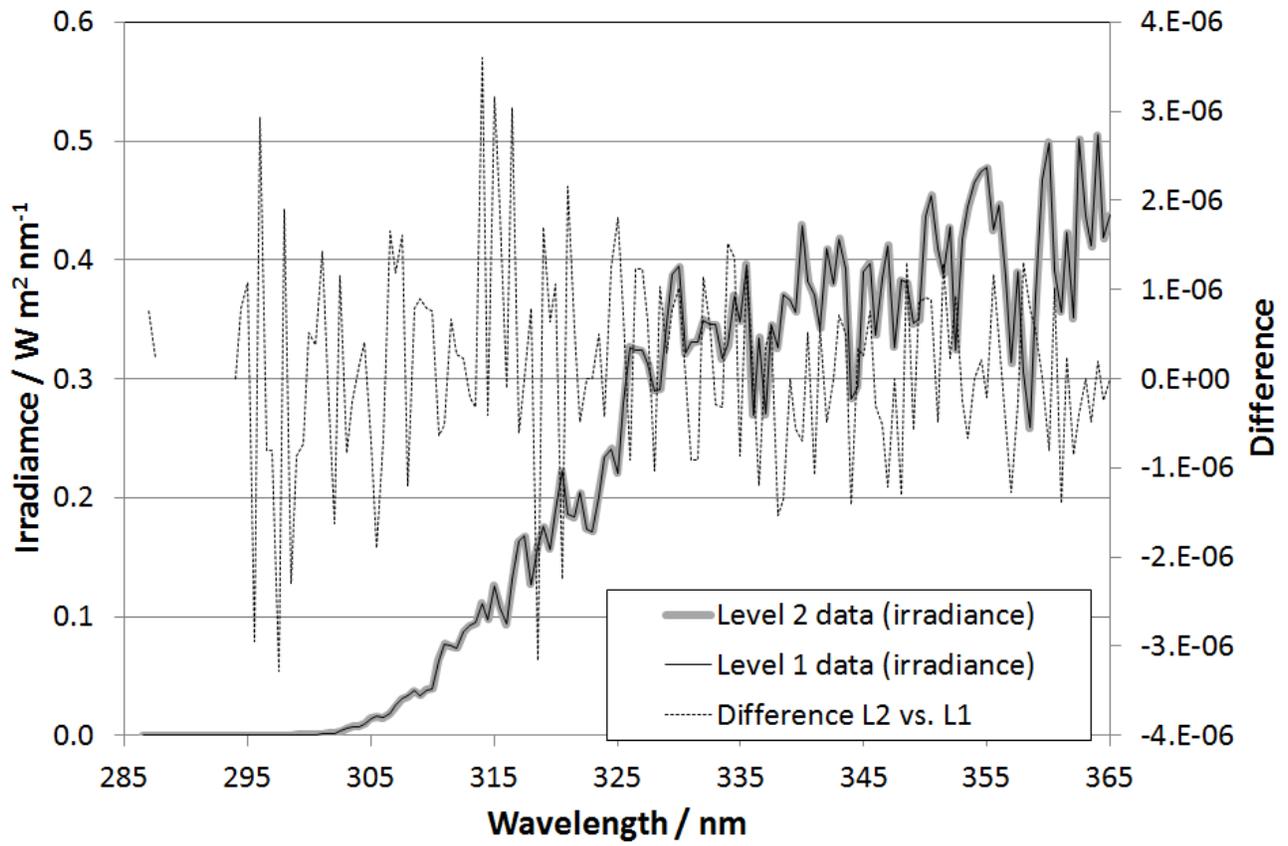


Figure 6. Schematic presentation of Brewer NRT data processing.

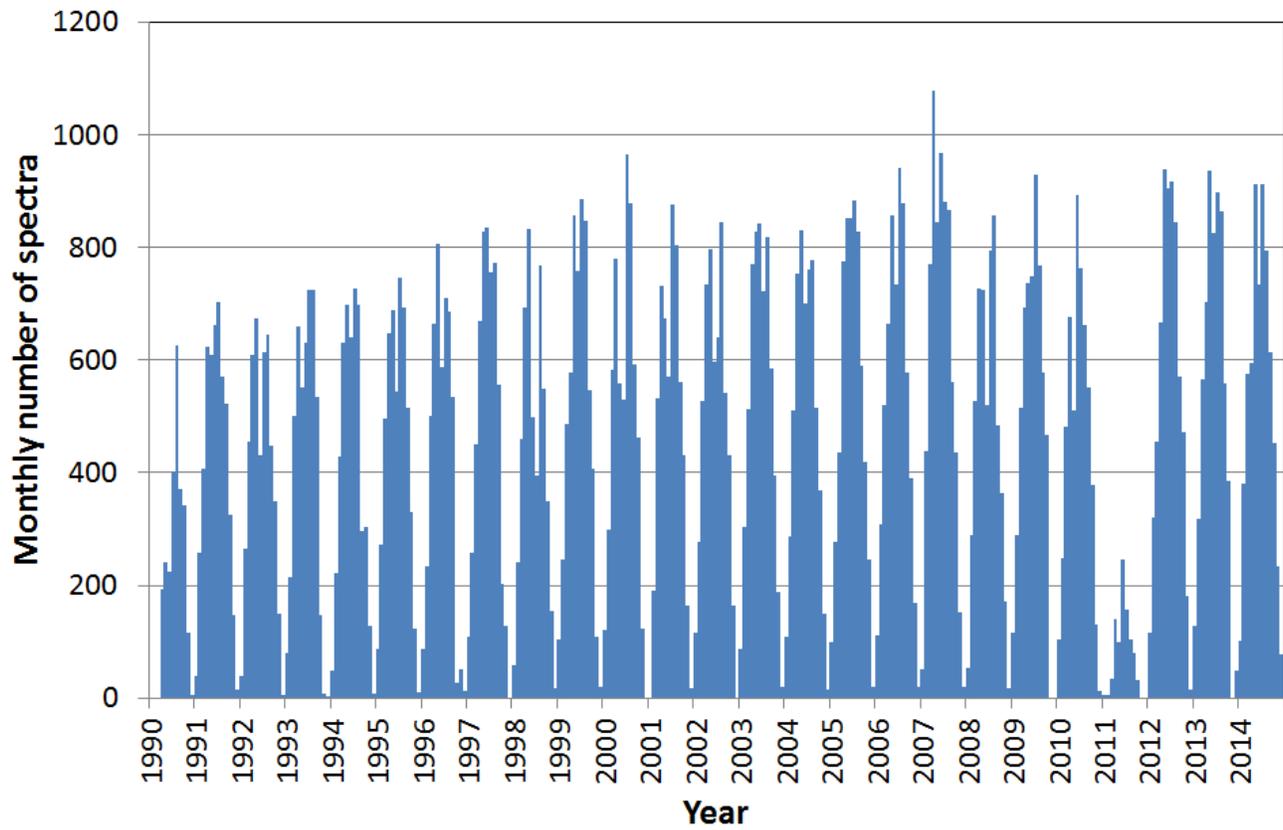


645

646

647 Figure 7. Examples of the Level 1 and Level 2 Brewer irradiances spectra.

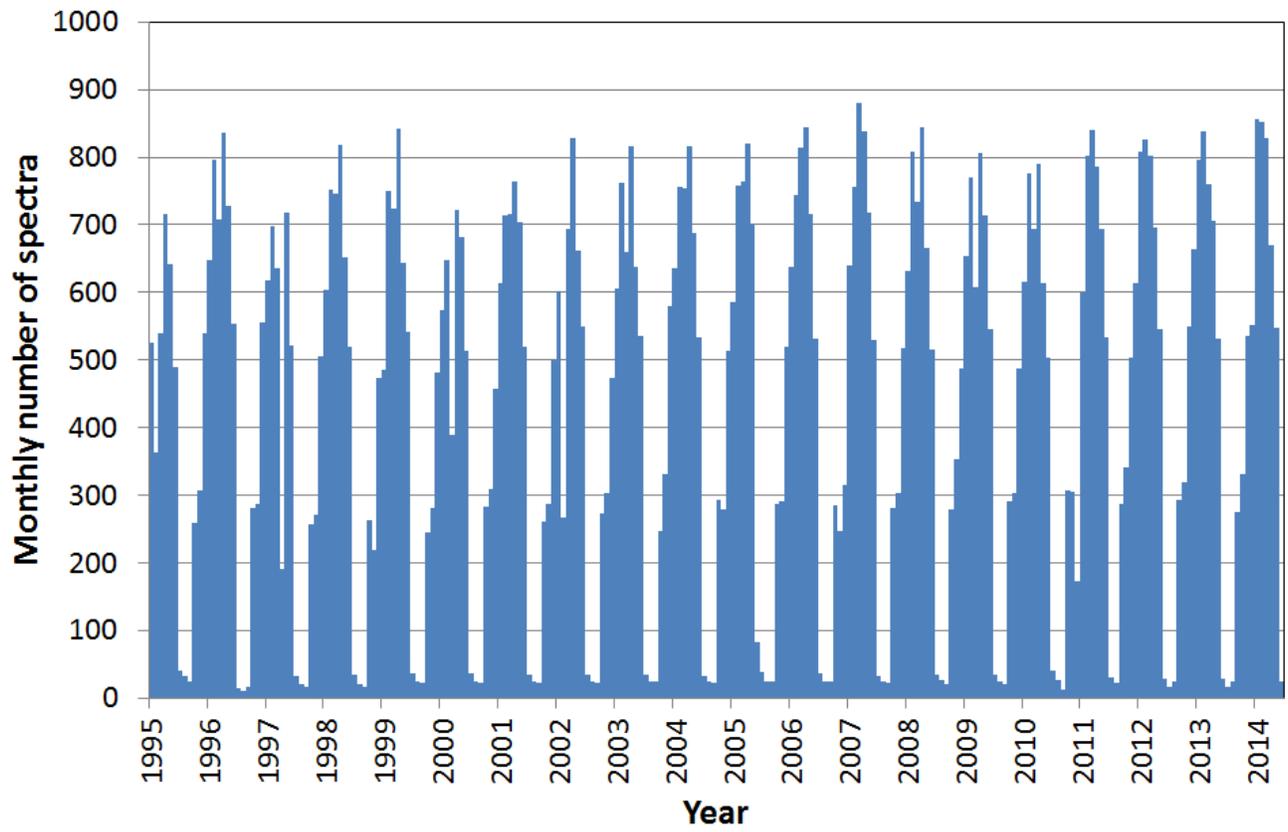
648



649

650 Figure 8a. Number of monthly UV irradiance spectra submitted to the EUVDB measured by the
 651 Brewer #037 during 1990-2014.

652



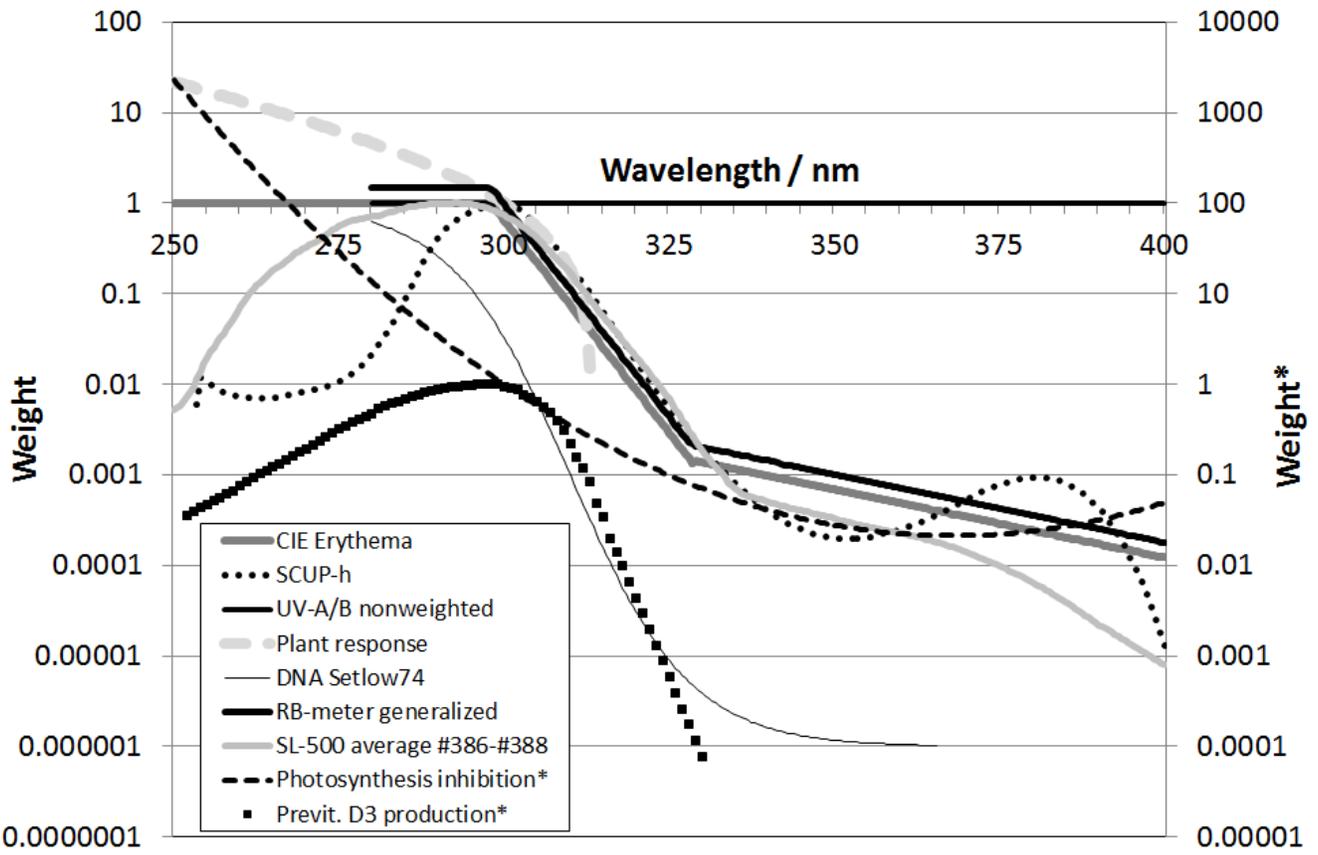
653

654

655

656 Figure 8b. Number of monthly UV irradiance spectra submitted to the EUVDB measured by the
 657 Brewer #107 during 1995-2014.

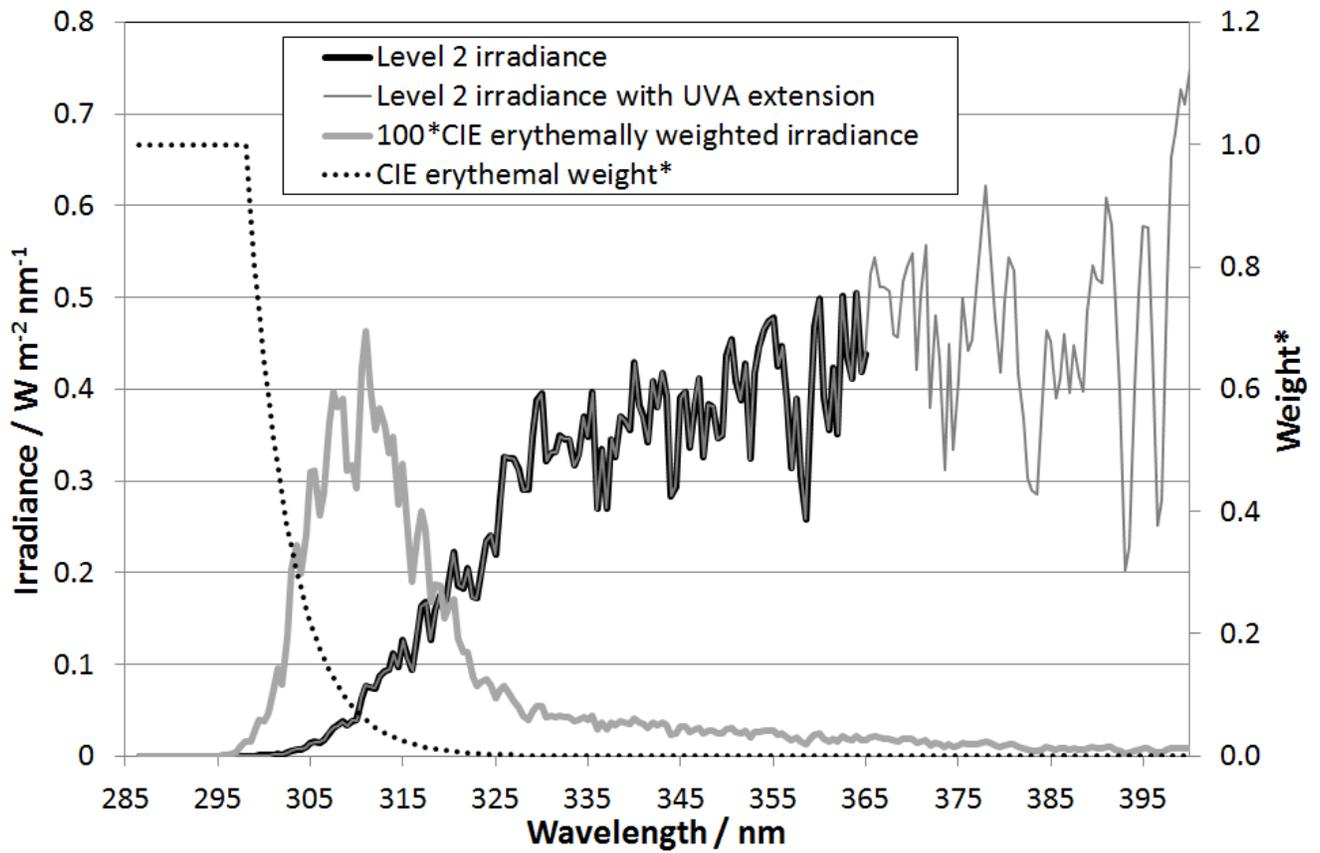
658



659

660 Figure 9. Action spectra used routinely in the processing to derive weighted dose rates from Brewer
 661 UV measurements.

662



663

664 Figure 10. Example of the UVA extension to a spectrum and calculation of the corresponding dose rate
 665 weighted with the CIE erythemal action spectrum.

666

667 Table 1. Information on the case UV scan of Jokioinen Brewer #107 selected for graphical
668 demonstration on the phases in the data flow.

Date	20 May 2007
Time of scan	13:20:07 – 13:25:18 UTC
SZA during the scan	51.3° – 51.4°
Total ozone column	325 DU
Visibility	30 km

669

670

671 Table 2. Action spectra used routinely in the processing to derive weighted dose rates from Brewer UV
672 measurements.

Action spectrum	Reference
Erythema (CIE weighting function)	McKinlay et al. 1987
Skin cancer in mice corrected for human skin, 299 nm normalization ()	de Gruijl, F.R. and J.C. van der Leun, 1994
UVB non weighted 290-320 nm	-
UVB non weighted 290-315 nm	-
UVA non weighted 315-400 nm	-
Generalized plant response (normalised at 300 nm) (Caldwell et al. 1986)	Caldwell et al. 1986
DNA damage (Setlow 1974)	Setlow 1974
Photosynthesis inhibition (Mitchell 1990)	Mitchell 1990
Previtamin D	CIE 2006

673
